

# 2025 COMMUNITY SPACE WEATHER MODELING AND DATA ASSIMILATION WORKSHOP



## Modeling Relativistic Electron Dropout in the Outer Radiation Belt During the 31 December 2016 Storm

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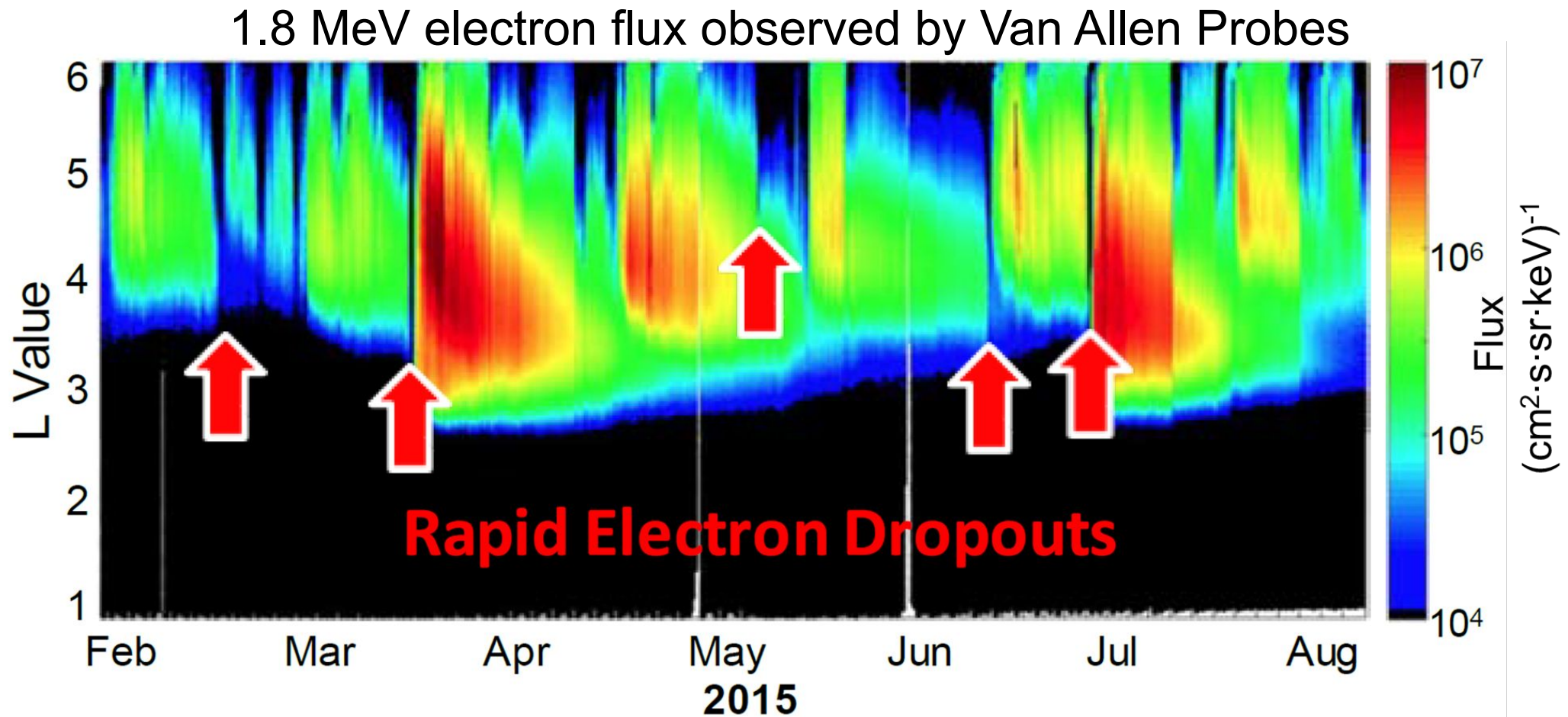
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Scan for full paper

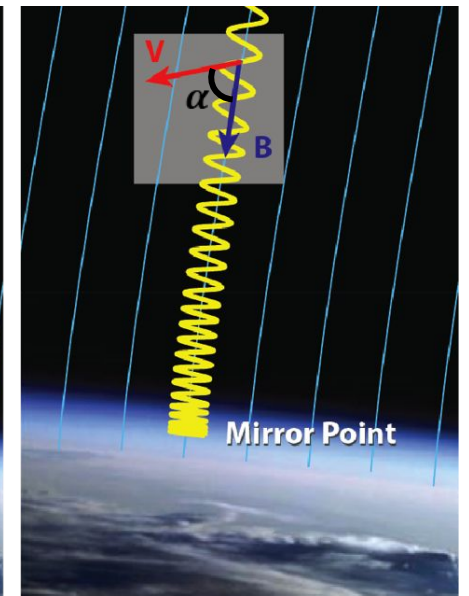
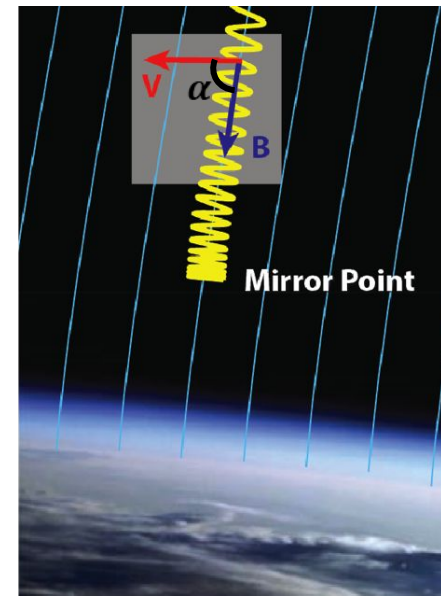
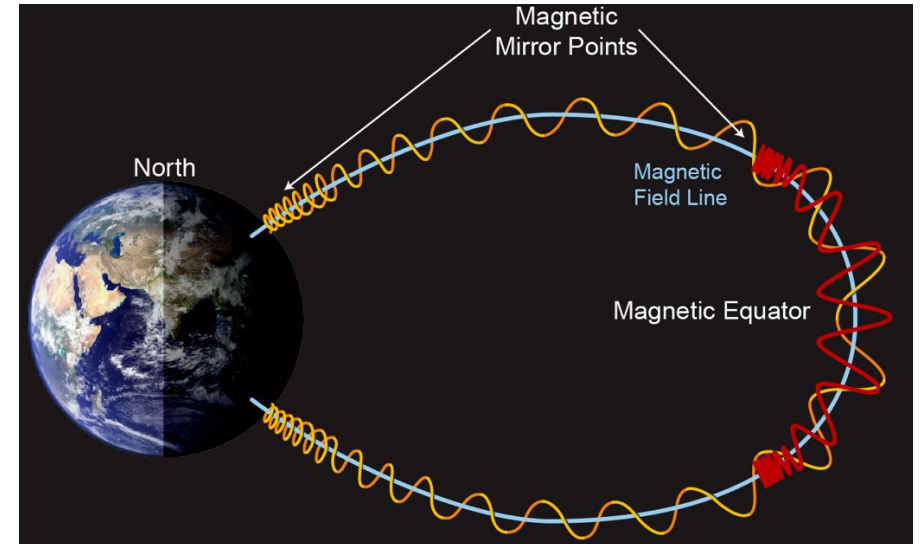
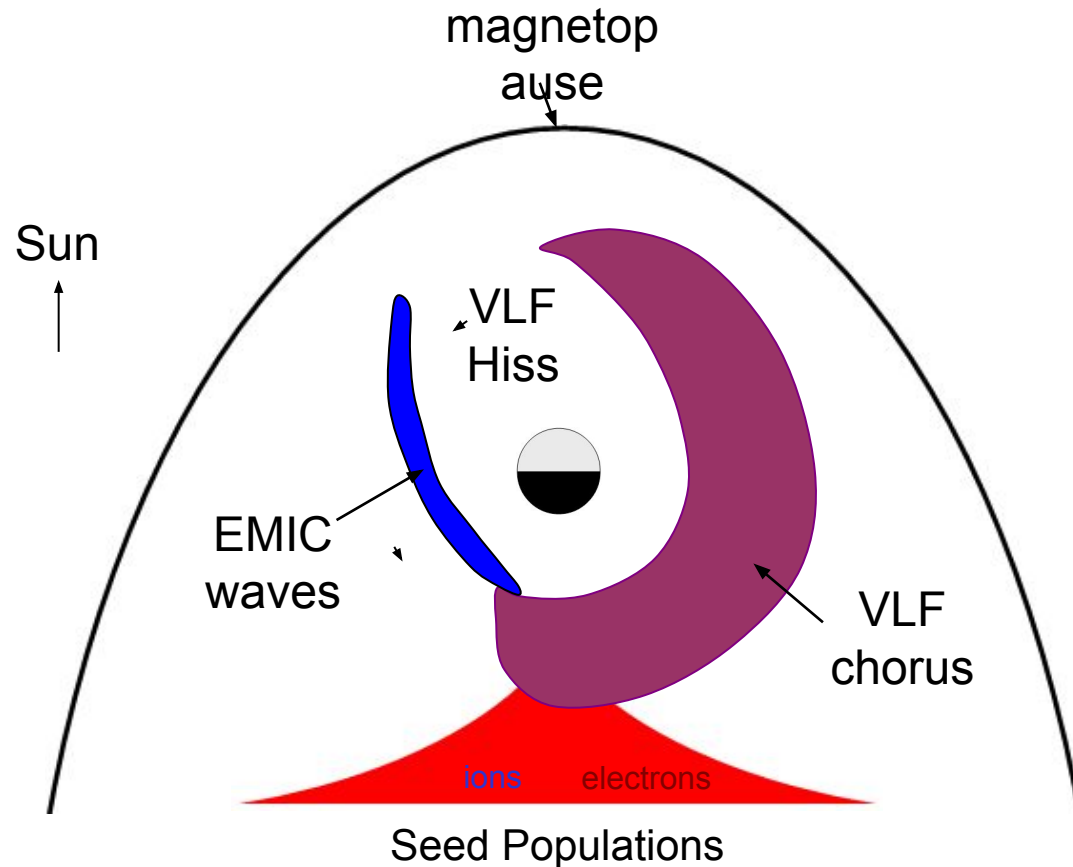
# Radiation Belt Electrons Loss



The fast dropout of the relativistic radiation belt electrons is one of the most **compelling and outstanding** questions in radiation belt studies.

# Radiation Belt Electrons Loss Mechanisms

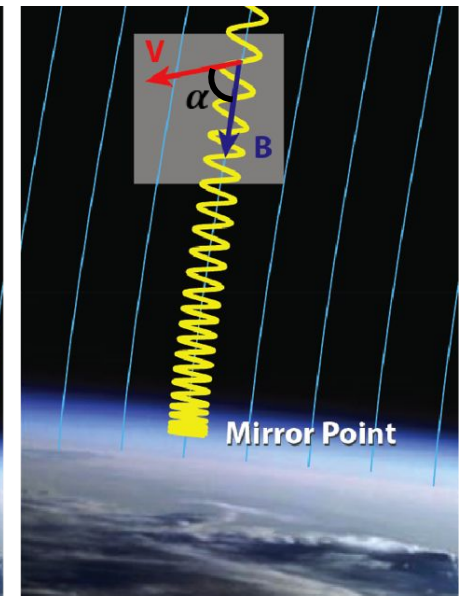
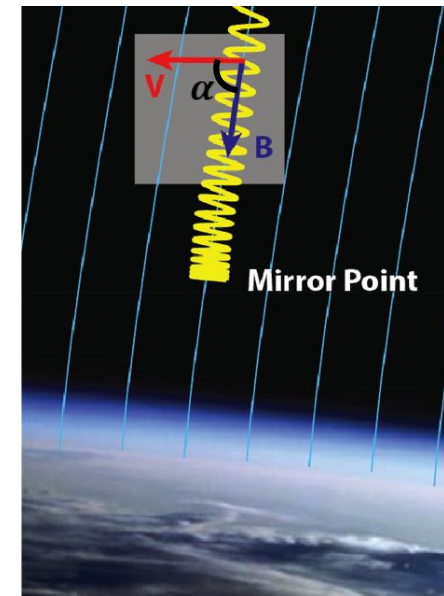
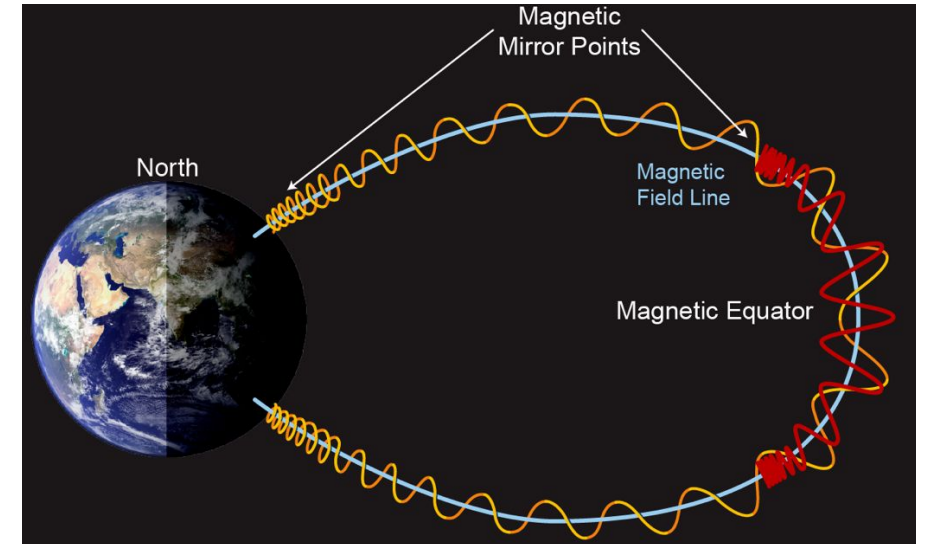
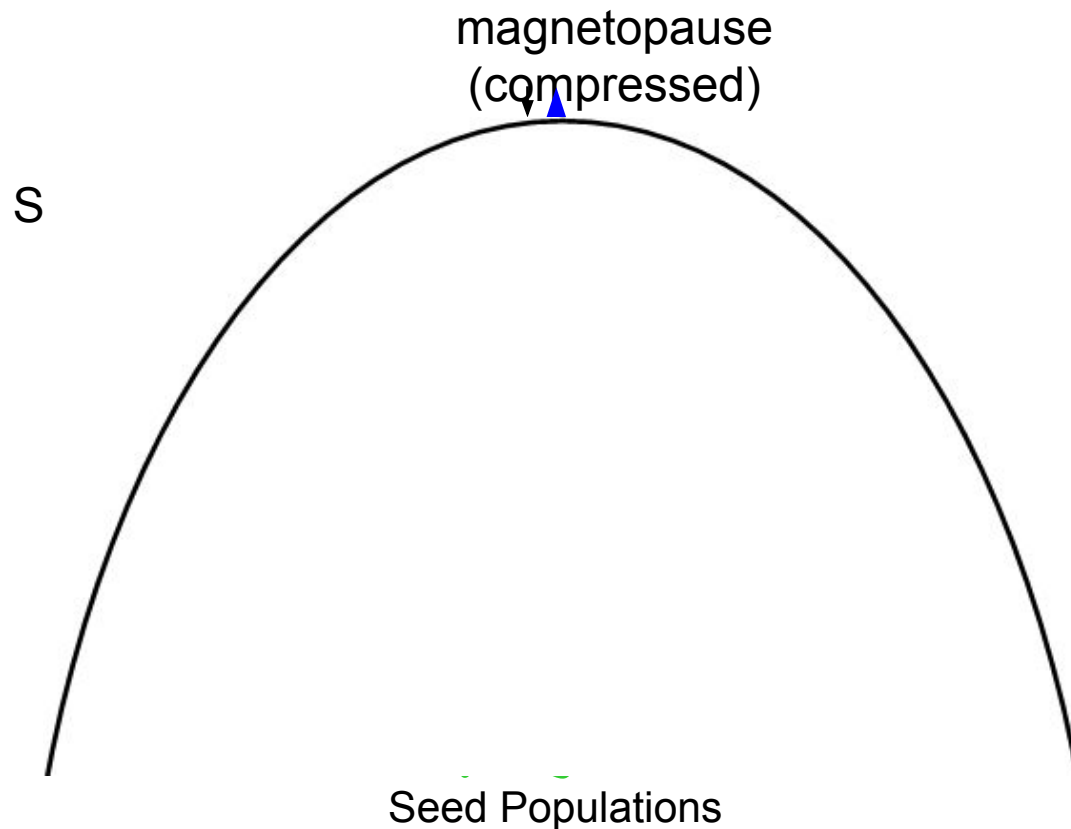
- **Precipitation loss:** pitch angle scattering by, e.g., VLF, EMIC waves, or field line curvature



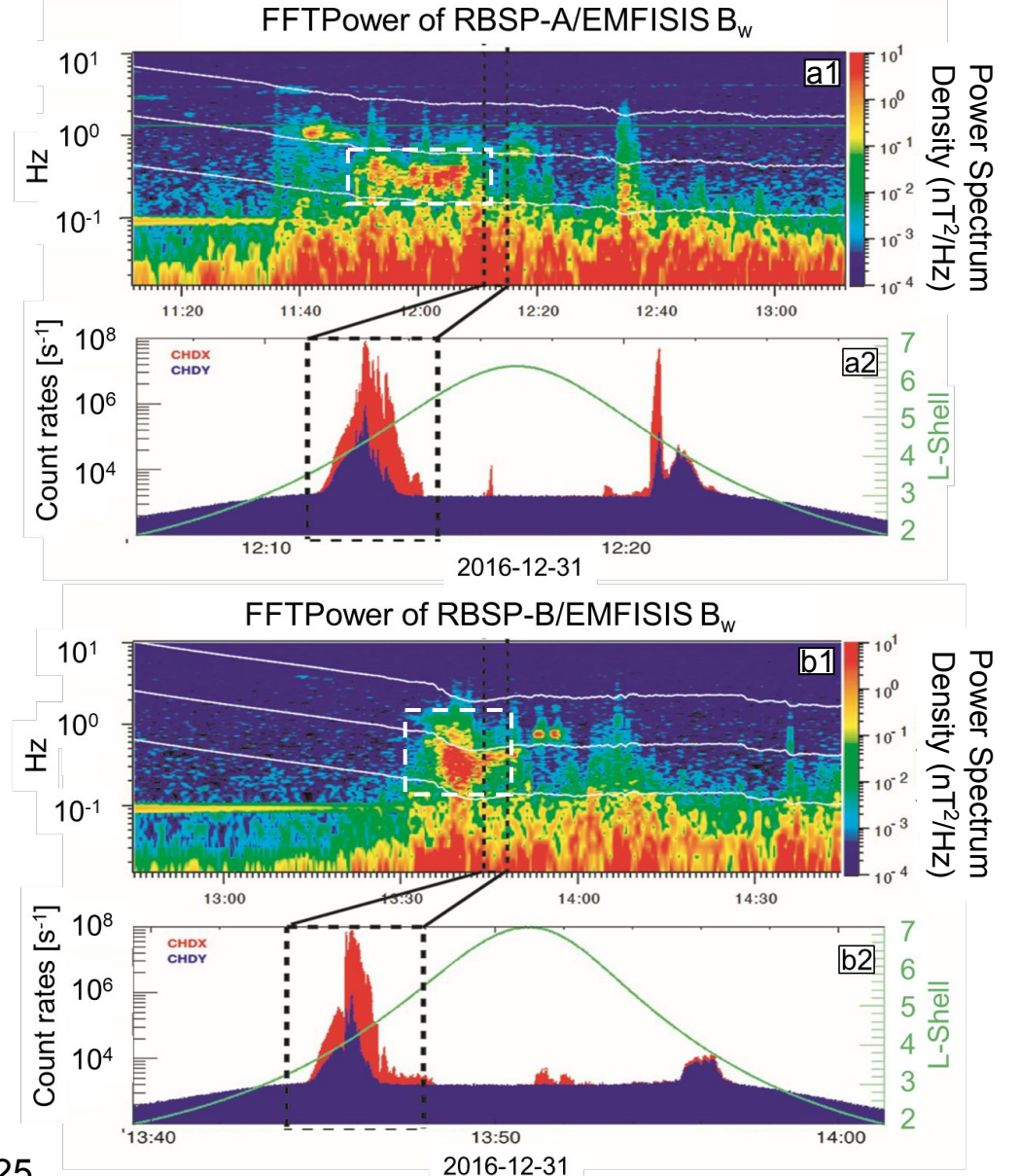
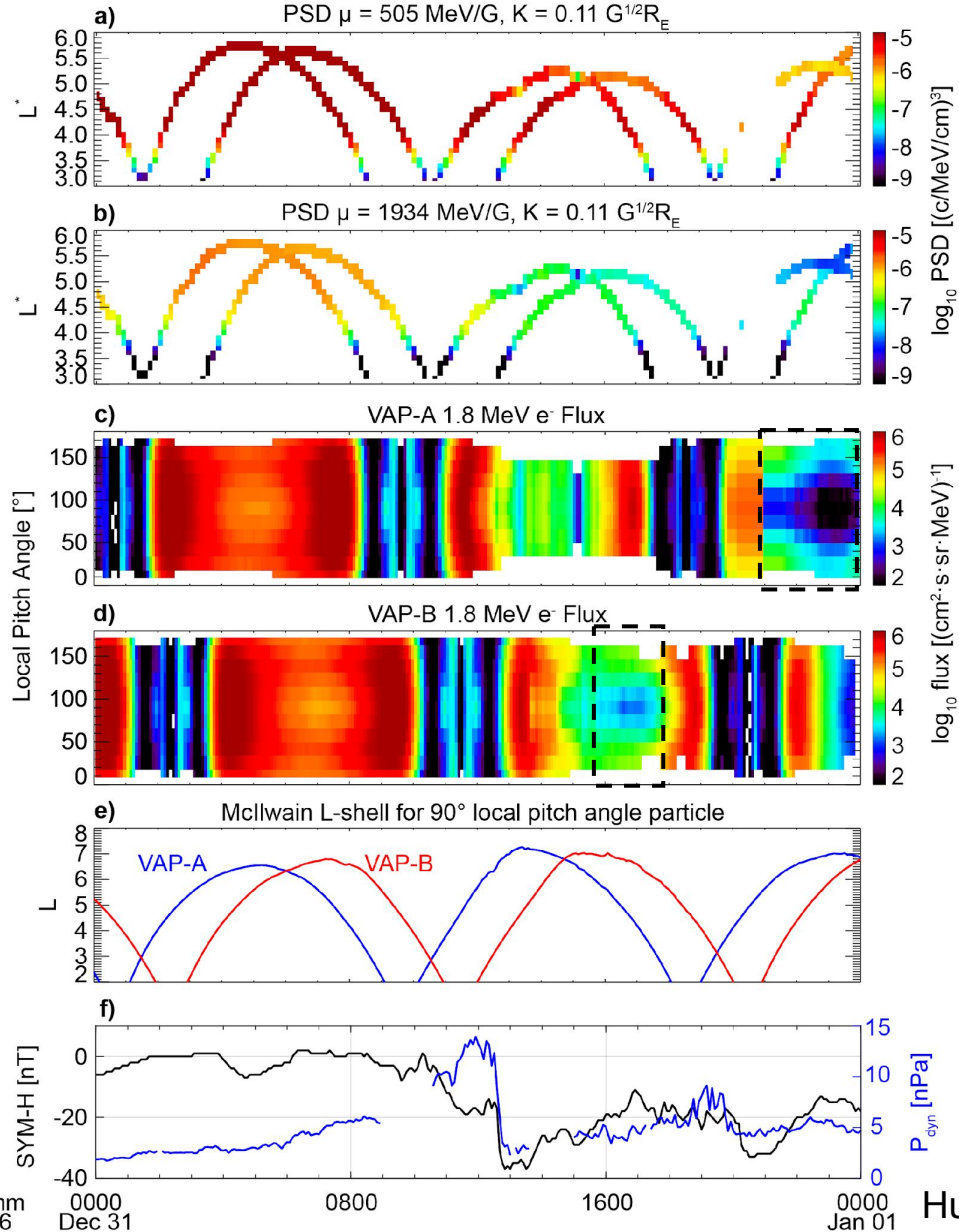


# Radiation Belt Electrons Loss Mechanisms

- **Precipitation loss**: pitch angle scattering by, e.g., **VLF, EMIC waves**, or **field line curvature**
- **Magnetopause shadowing**: combined with outward radial transport by **ULF waves**



# Event Overview



Huang et al., 2025

Bruno et al., 2022

# Model Description

- 2D diffusion model (radial and pitch angle)

Radial diffusion by **ULF** waves (Murphy et al., 2023)

$$\log_{10}(D_{L^*L^*}) = c + a_1 L^* + a_2 \text{Sym}H + a_3 B_z + a_4 V + a_5 P_{dyn}$$

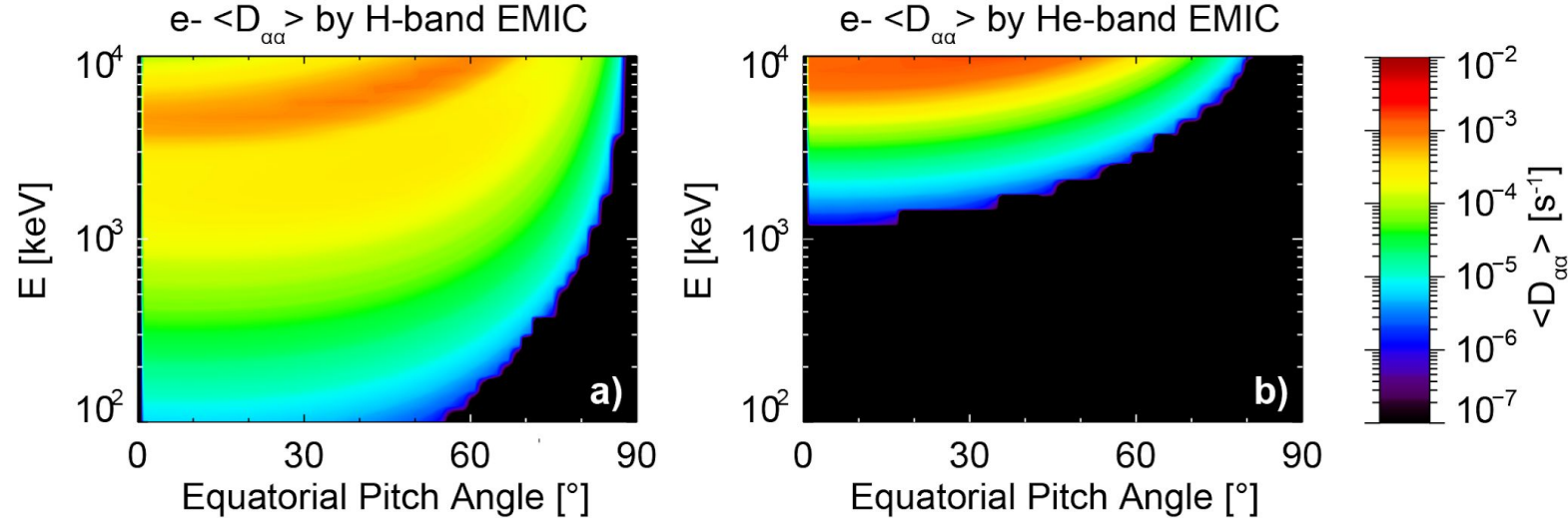
$$\frac{\partial f}{\partial t} = L^{*2} \frac{\partial}{\partial L^*} \left( \frac{D_{L^*L^*}}{L^{*2}} \frac{\partial f}{\partial L^*} \right) + \frac{1}{G} \frac{\partial}{\partial \alpha} \left( G \langle D_{\alpha\alpha} \rangle \frac{\partial f}{\partial \alpha} \right) - \frac{f}{\tau}$$
$$G = T(\alpha) \sin(2\alpha), T(\alpha) = 1.38 - 0.32 (\sin(\alpha) + \sqrt{\sin(\alpha)})$$

Pitch angle diffusion by **EMIC** waves (based on Zhang et al., 2016), or by field line curvature (**FLC**) scattering (Young et al., 2008)

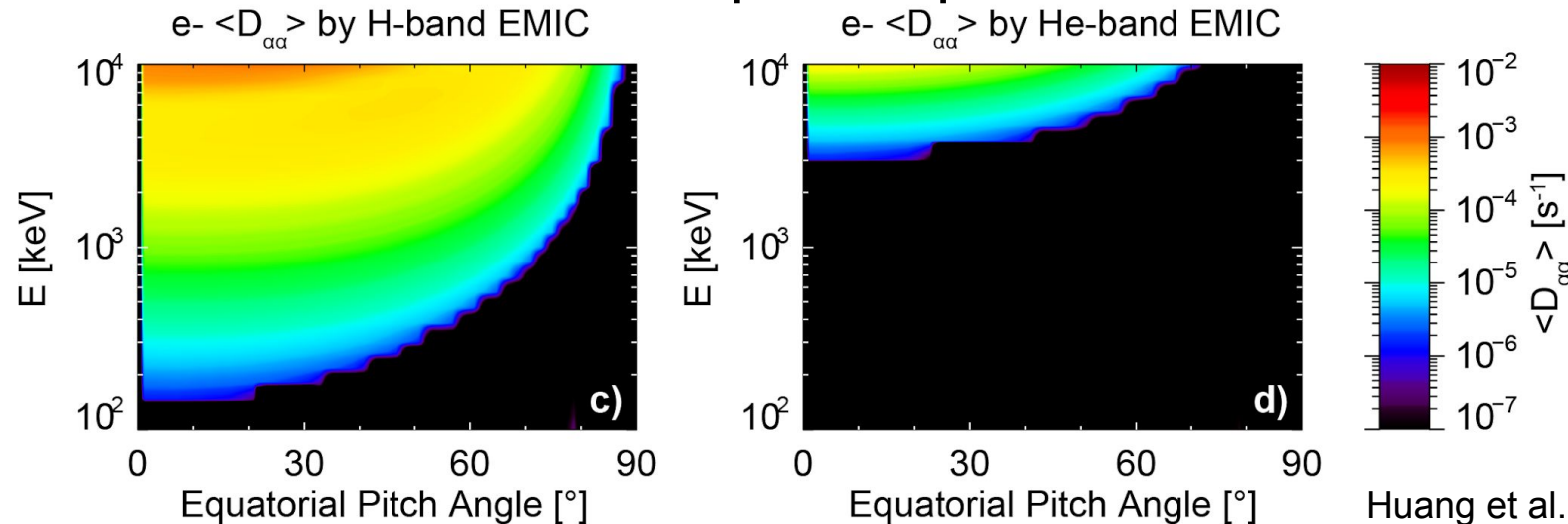
Electron lifetimes: on the order of electron drift periods outside the last closed drift shell (**LCDS**) (Huang et al., 2023) or inside the drift loss cone

# $\langle D_{\alpha\alpha} \rangle$ by EMIC

L = 4.0, Inside plasmapause



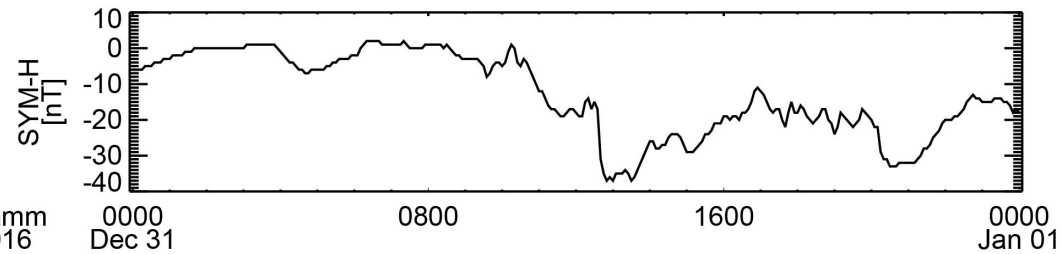
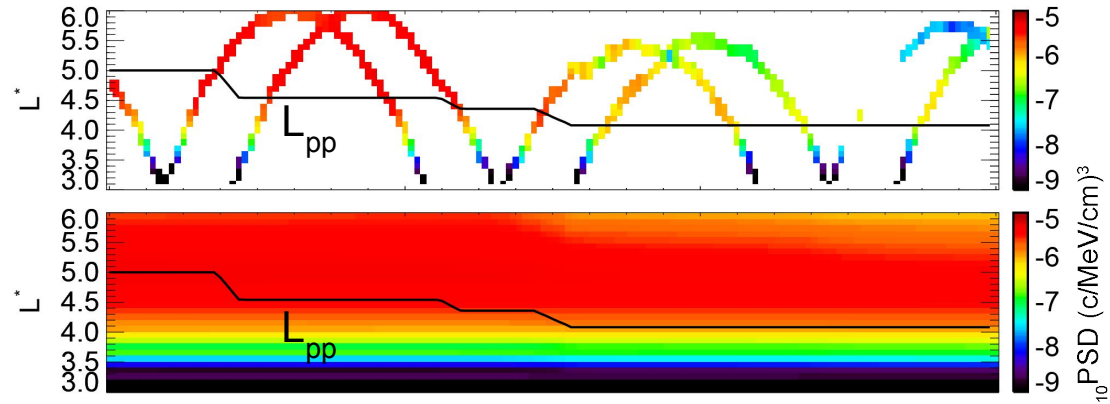
L = 5.0, Outside plasmapause





# Simulation Results

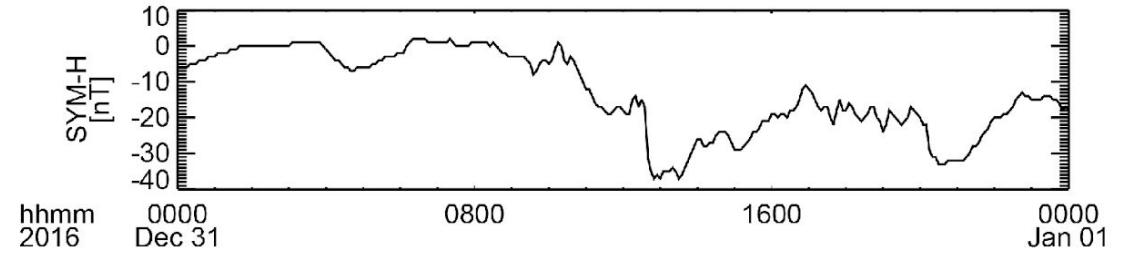
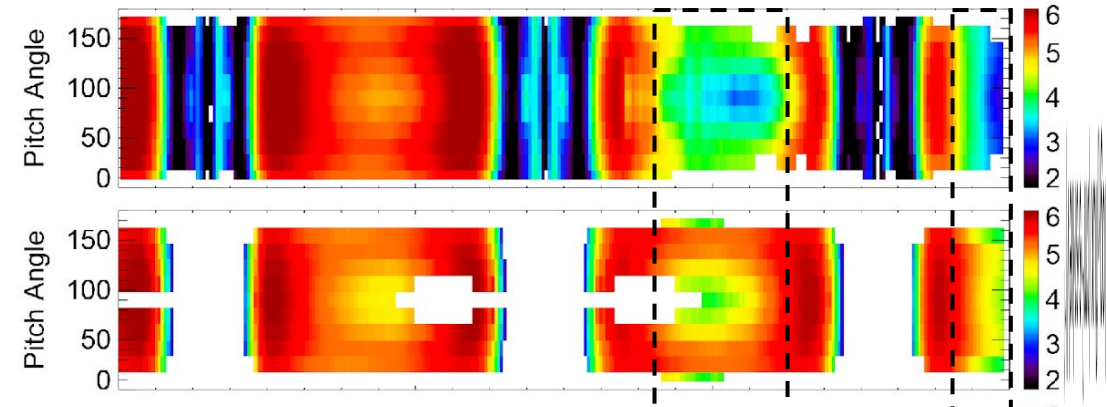
e- PSD:  $\mu = 1934 \text{ MeV/G}$ ,  $K = 0.03 R_E G^{1/2}$



$$L_{pp} = 5.6 - 0.46 K p_{max}$$

Carpenter and Anderson, 1992

VAPB 1.8 MeV e- Flux PAD

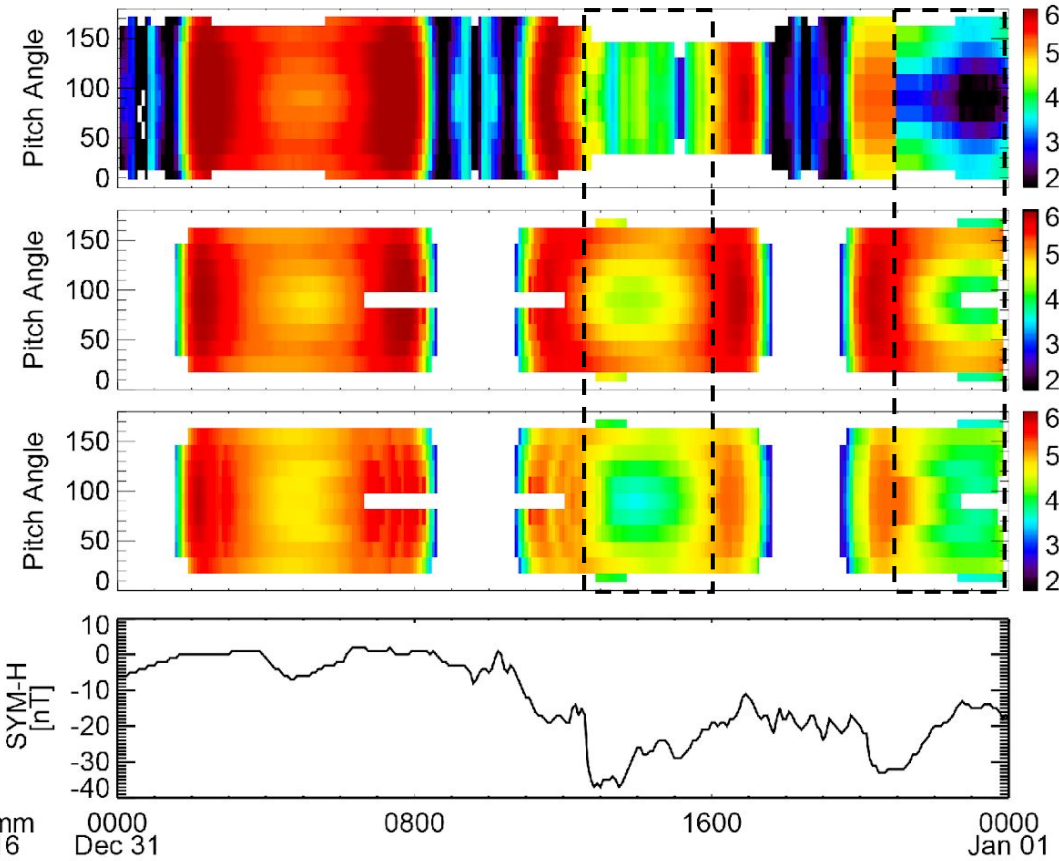


Huang et al., 2025

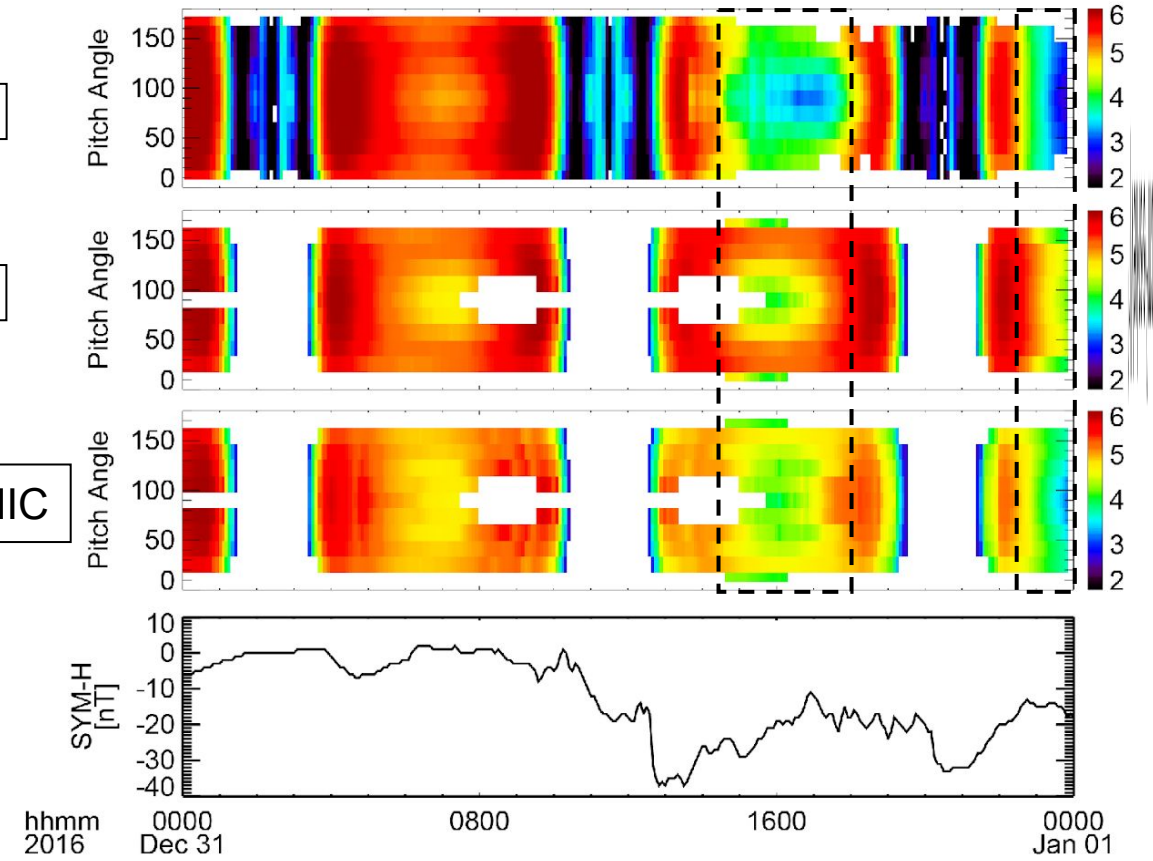


# Simulation Results

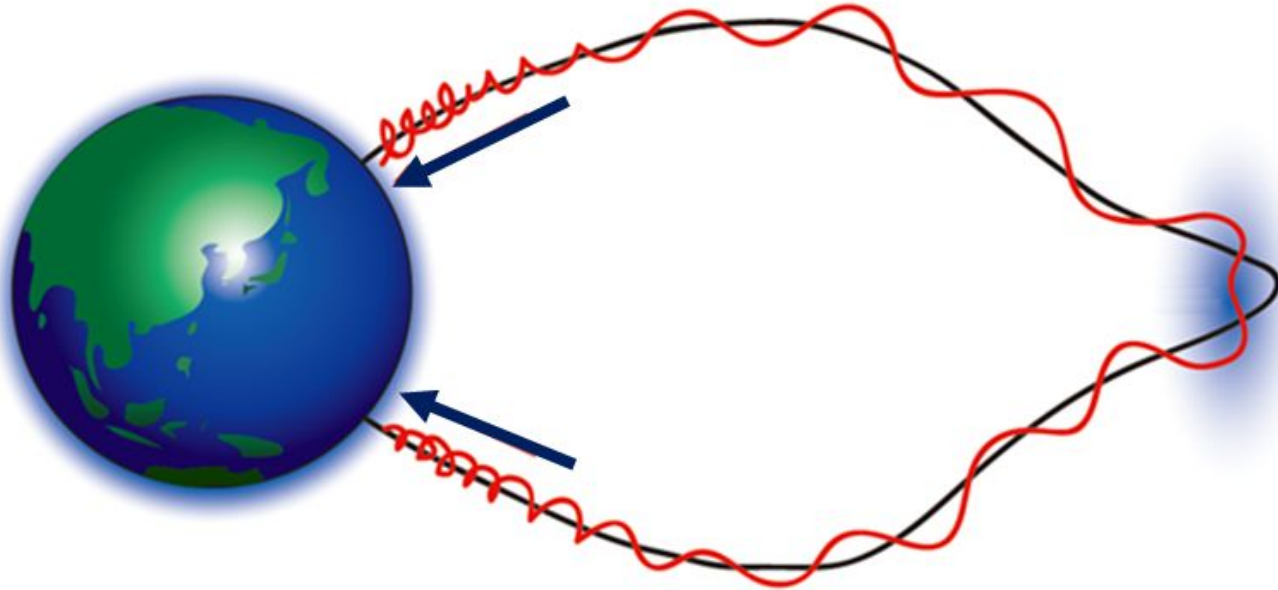
VAPA 1.8 MeV e- Flux PAD



VAPB 1.8 MeV e- Flux PAD



# $\langle D_{\alpha\alpha} \rangle$ by FLC



Field line curvature scattering:  $\epsilon = R_g/R_c > 0.1$

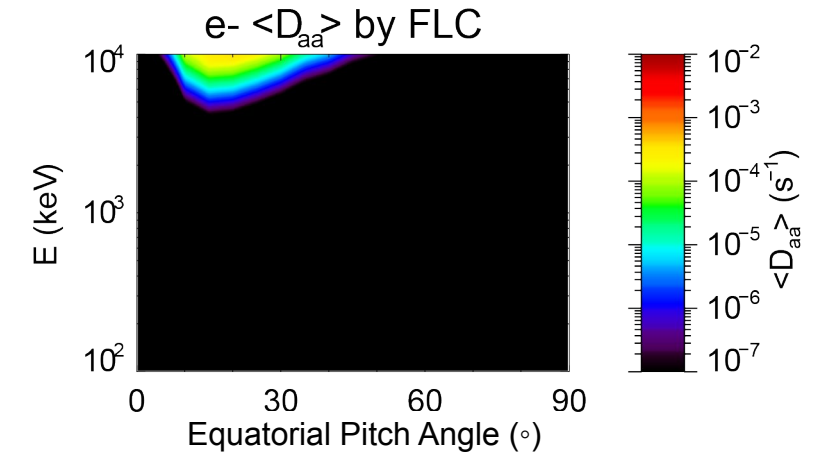
For  $\epsilon < 0.584$ :

$D_{\alpha\alpha}(\epsilon, \tau_B, \alpha_{eq}, \text{field geometry})$  (Young et al., 2008)

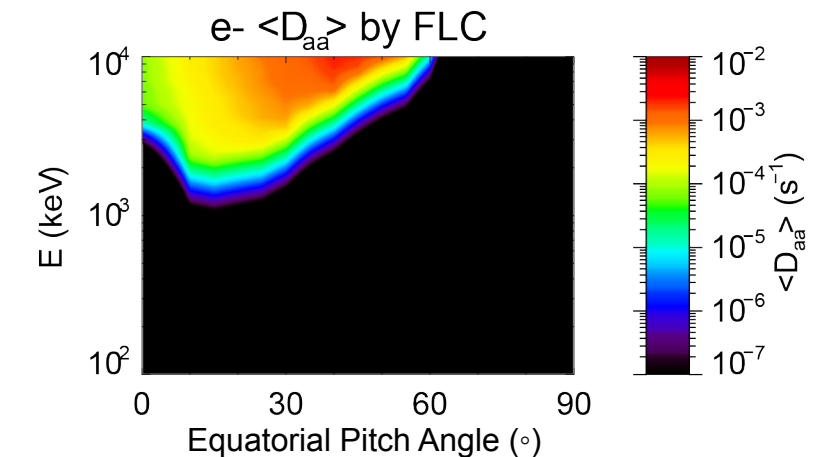
For  $\epsilon \geq 0.584$ :

Strong diffusion:  $D_{SD} = 2\alpha_{BLC}^2/\tau_B$  (Kennel, 1969)

$L^* = 5.0$ , TS04, 12/31/2016 13UT

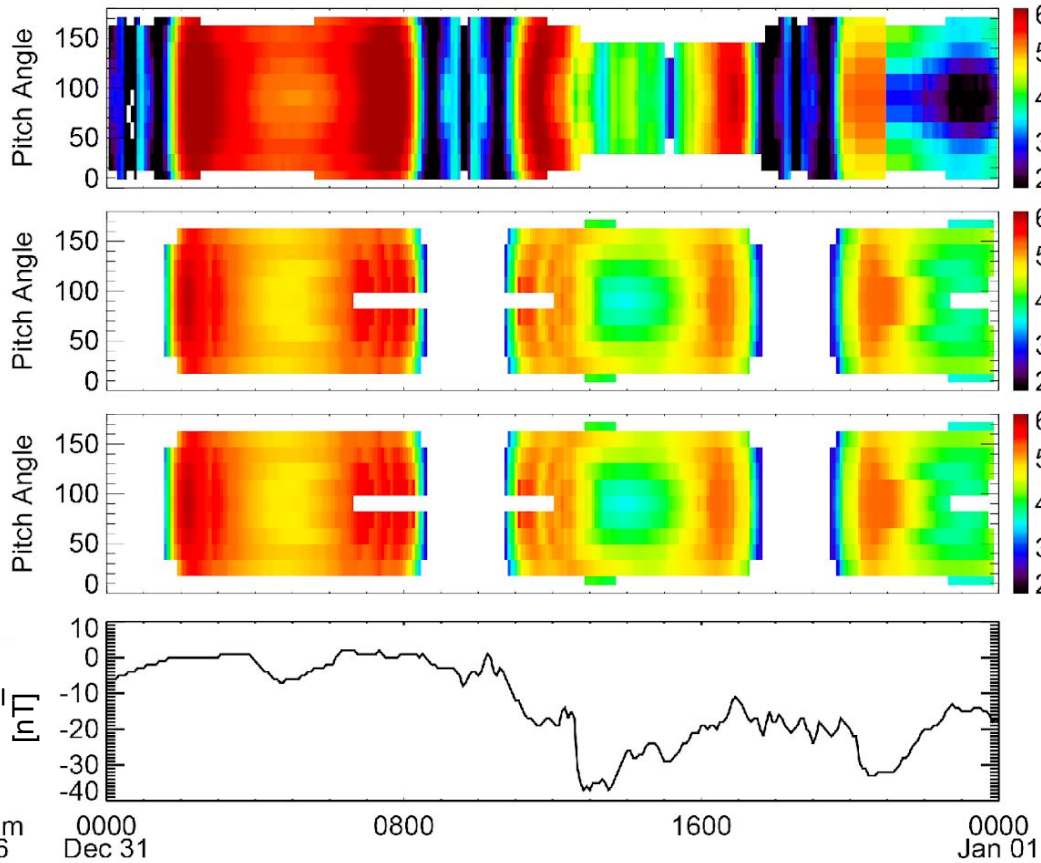


$L^* = 5.5$ , TS04, 12/31/2016 13UT

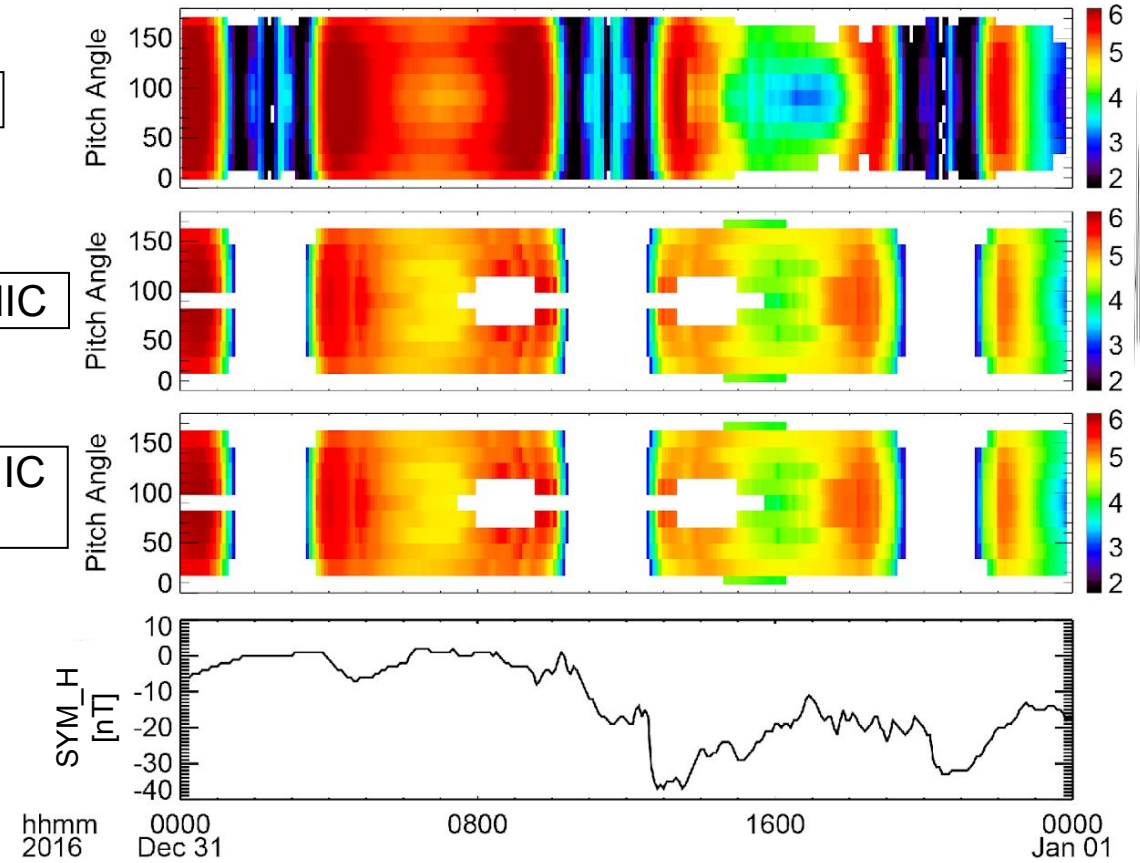


# Simulation Results

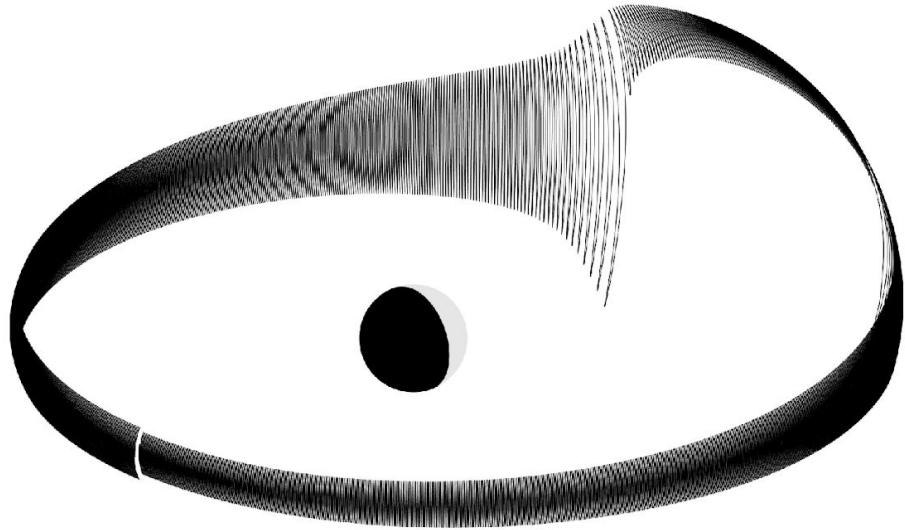
VAPA 1.8 MeV e- Flux PAD



VAPB 1.8 MeV e- Flux PAD

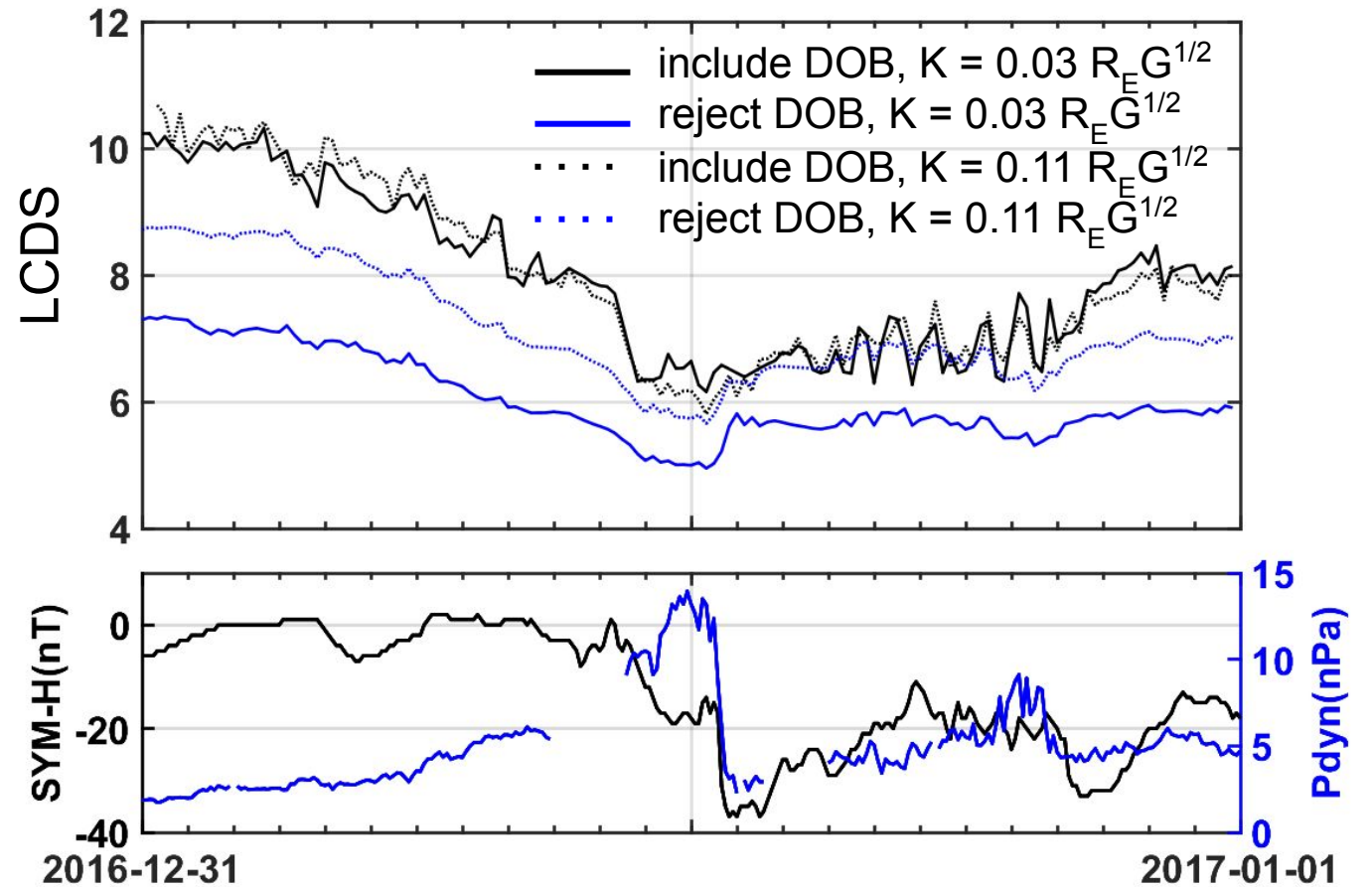


# Uncertainties in LCDS



e- Drift Orbit Bifurcation

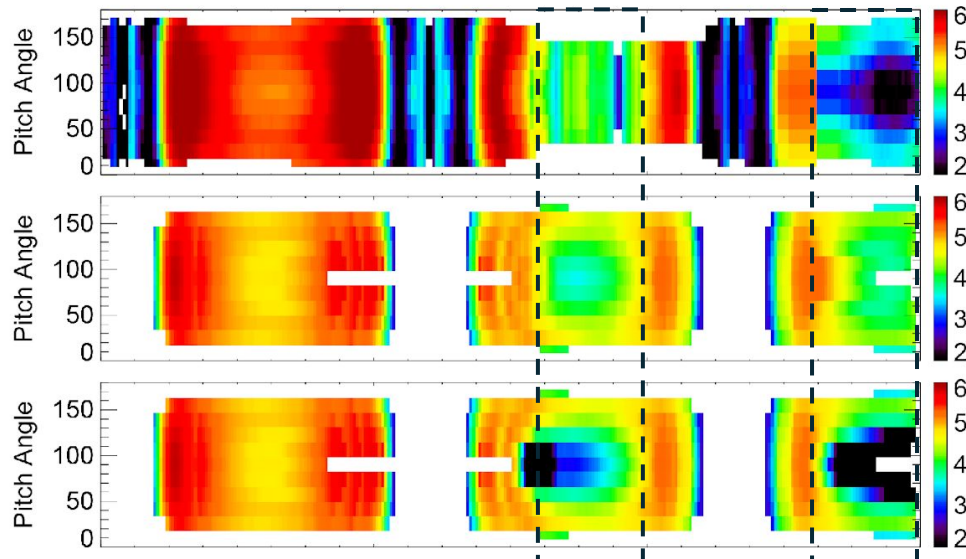
TS04



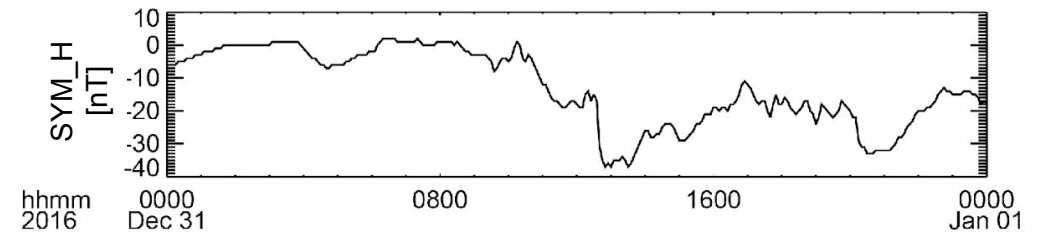
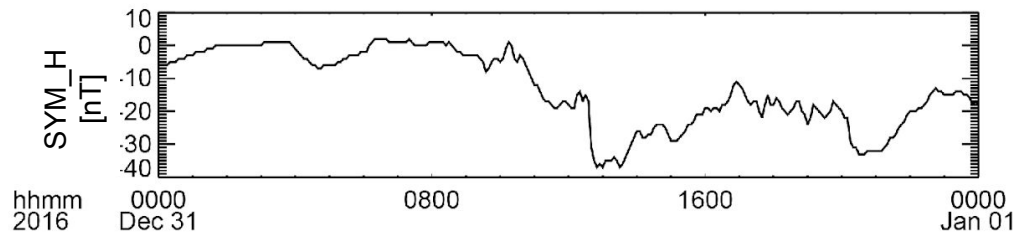
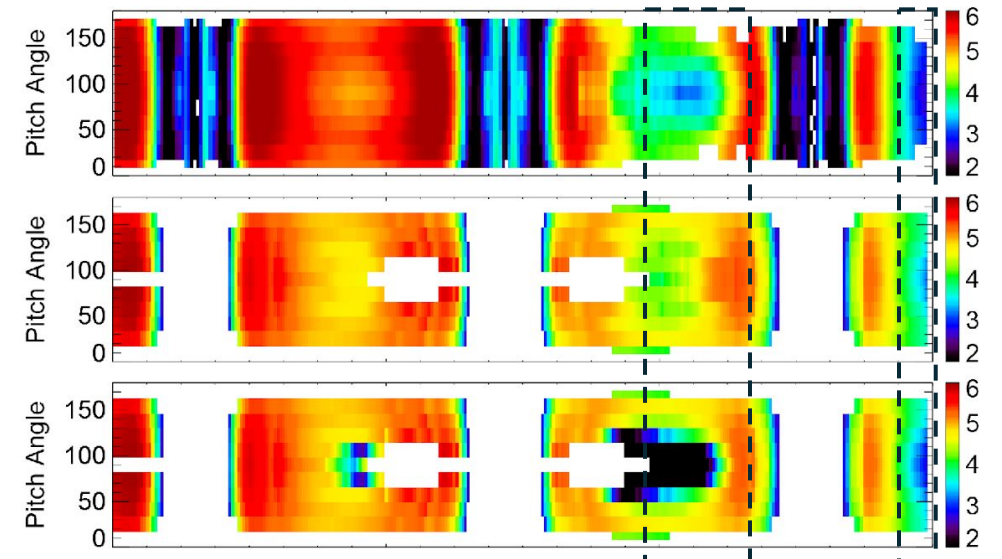


# Simulation Results

VAPA 1.8 MeV e- Flux PAD



VAPB 1.8 MeV e- Flux PAD



# Conclusions

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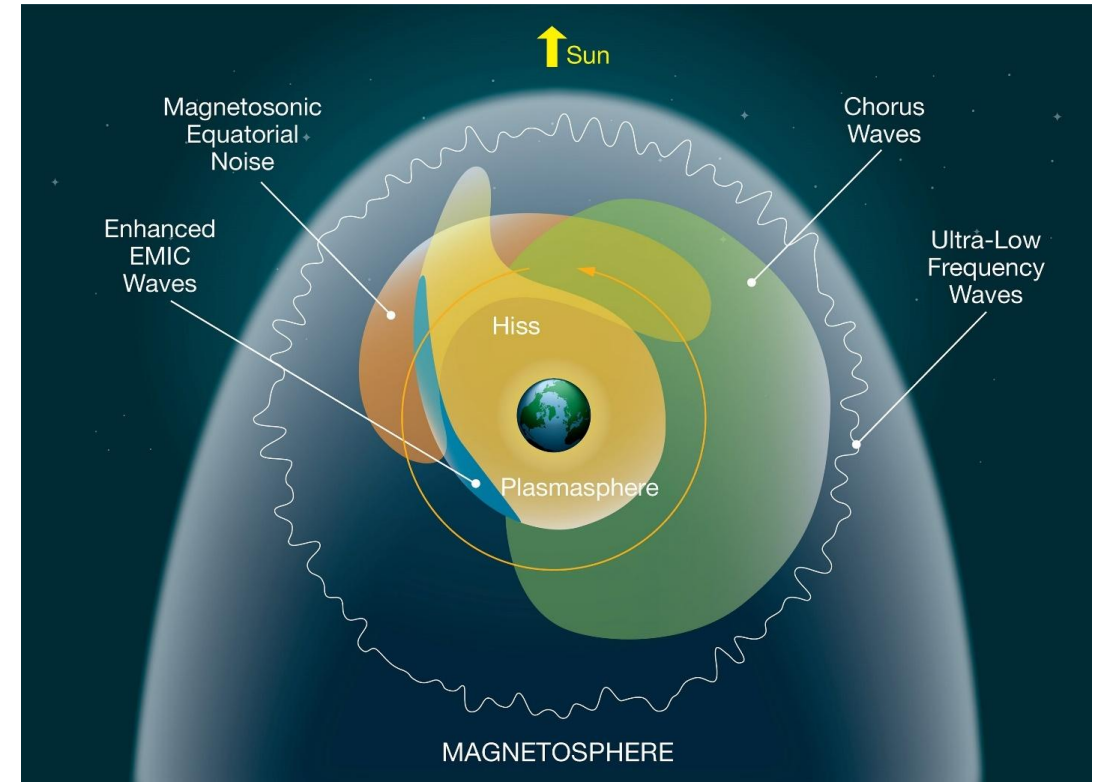
- 1. The dropout of relativistic outer radiation belt electrons mainly results from the **combined effects** of MPS and EMIC wave scattering. **MPS** dominates the dropout at **high L** and **high equatorial pitch angles**, while **EMIC** wave scattering is the primary mechanism at **low equatorial pitch angles** over **a wide L range** in the outer radiation belt.
- 2. The FLC effect is negligible in contributing to the observed electron dropout inside the LCDS.
- 3. The diffusion model requires physical quantification of MPS and more realistic and event-specific wave properties for EMIC wave scattering to better reproduce the observed electron dropout.



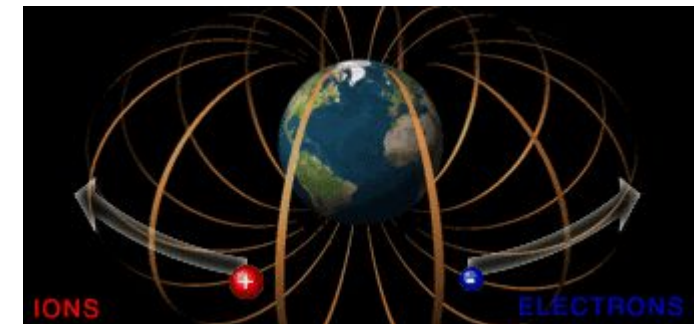
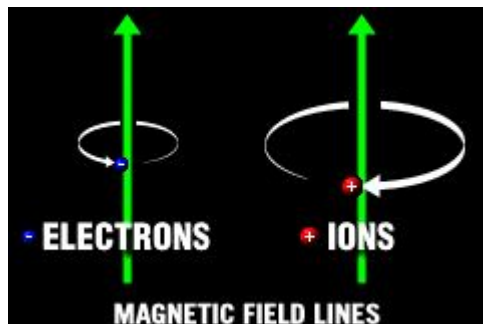
# **Backup Slides**

# Inner Magnetosphere: Waves

- A variety of **electromagnetic waves** exist in near-Earth space environment.
  - ULF waves, Chorus, Hiss, EMIC, etc.
- These waves have different frequencies that can **resonate** with the characteristic motions of charged particles.
  - Drift resonance, bounce resonance, gyro-resonance



Credit: NASA





# EMIC wave (Zhang et al., 2016)

% ADOPTED PARAMETERS (EMIC waves):

% B\_w=1000.0 pT; MLT=21.0 ;  
lambda\_min= 0.0 deg, lambda\_max=40.0 deg;

% Density Model=C; Ne=  
0.1178923E+009m^-3;  
fpe/fce= 14.08; N\_res = -5 ~ +5;

% omega\_lc= 0.25 omega\_eq,  
omega\_uc= 0.99 omega\_eq;

% theta\_m= 0.0 deg, theta\_w= 0.1 deg,  
theta\_lc= 0.0 deg, theta\_uc= 0.2 deg.

