

Next-Generation MHD Modeling of Solar Wind Using Neural Operators

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UCAR
COMMUNITY
PROGRAMS



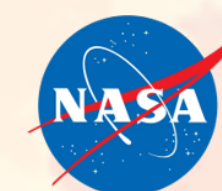
CPAESS



Enrico Camporeale
Thomas Berger



Arpit Shrivastav

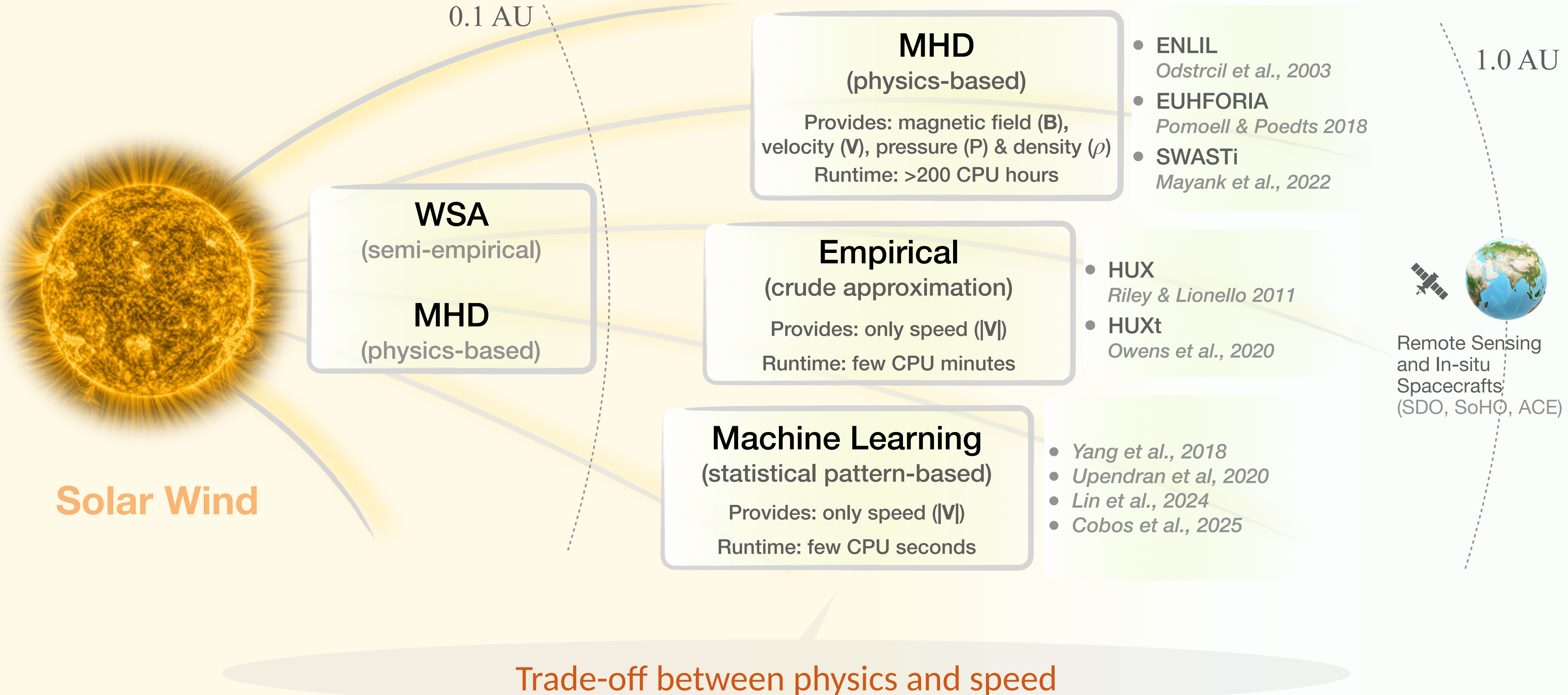


Nick Arge

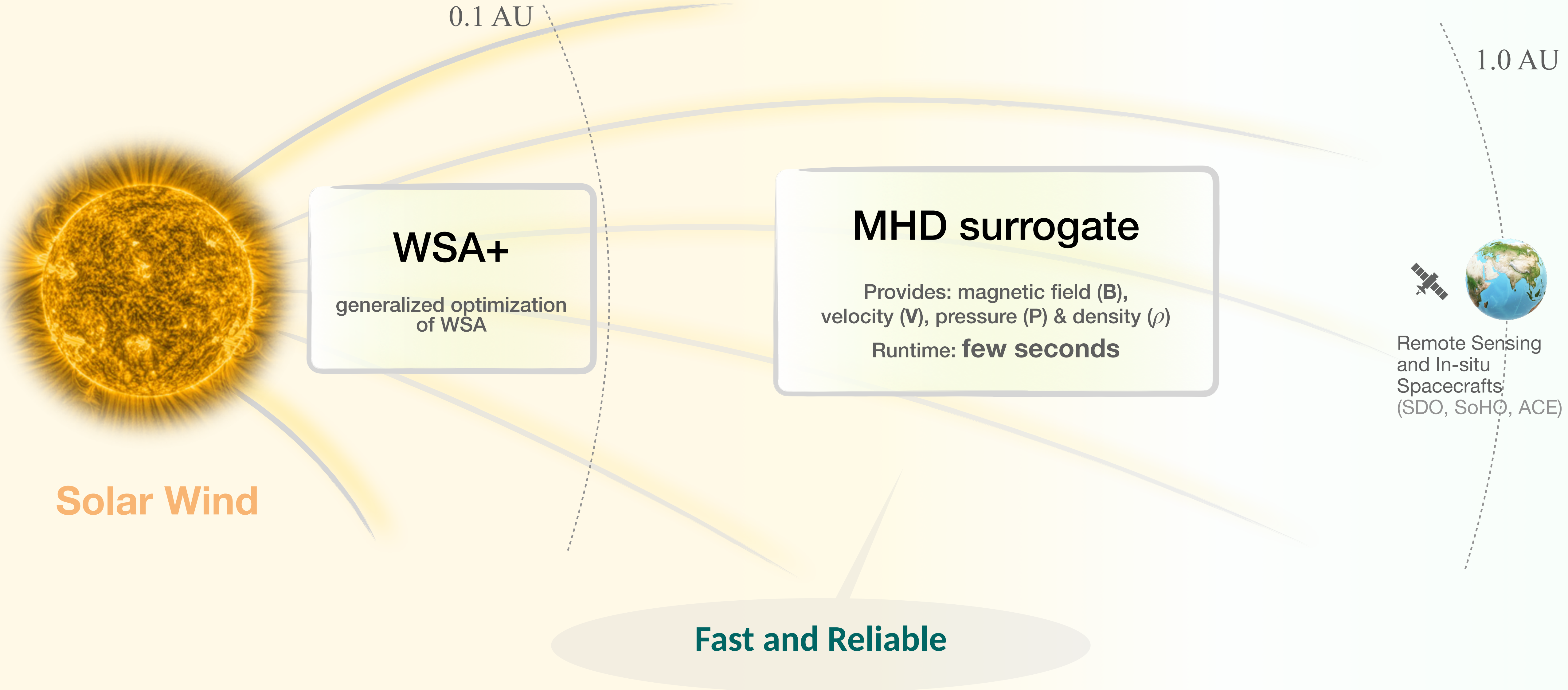


Zhenguang Huang
Gabor Toth

Current Modeling Approaches



Next Generation Approach



WSA – Solar Wind Relation

$$V = v_{\min} + \frac{v_{\max}}{(1 + f_s)^{\alpha}} \times \left(\left(1.0 - a1 \exp \left(\frac{-D}{w} \right)^{\beta} \right)^{a2} \right)$$

Magnetogram derived quantities:

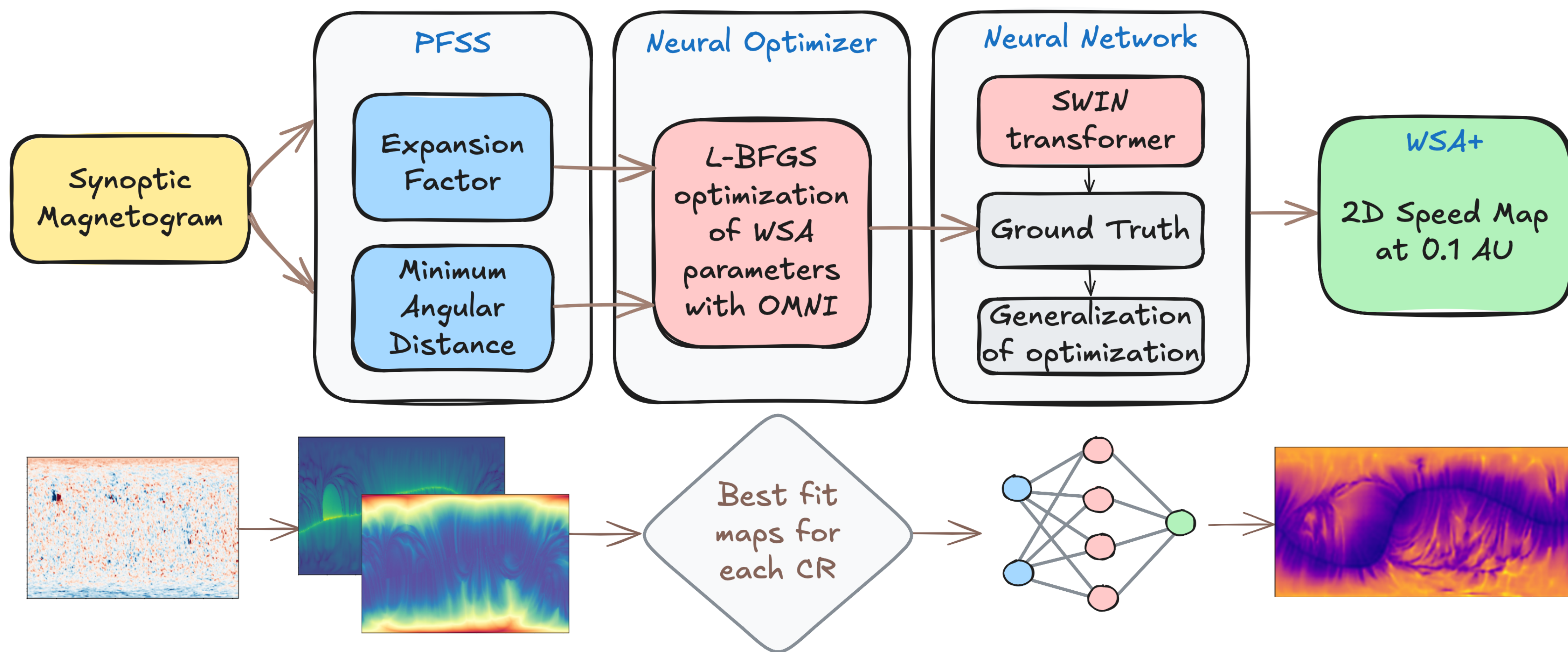
1. expansion factor
2. minimum angular distance from coronal hole boundary

$$v_{\max} = 750, v_{\min} = 250,$$

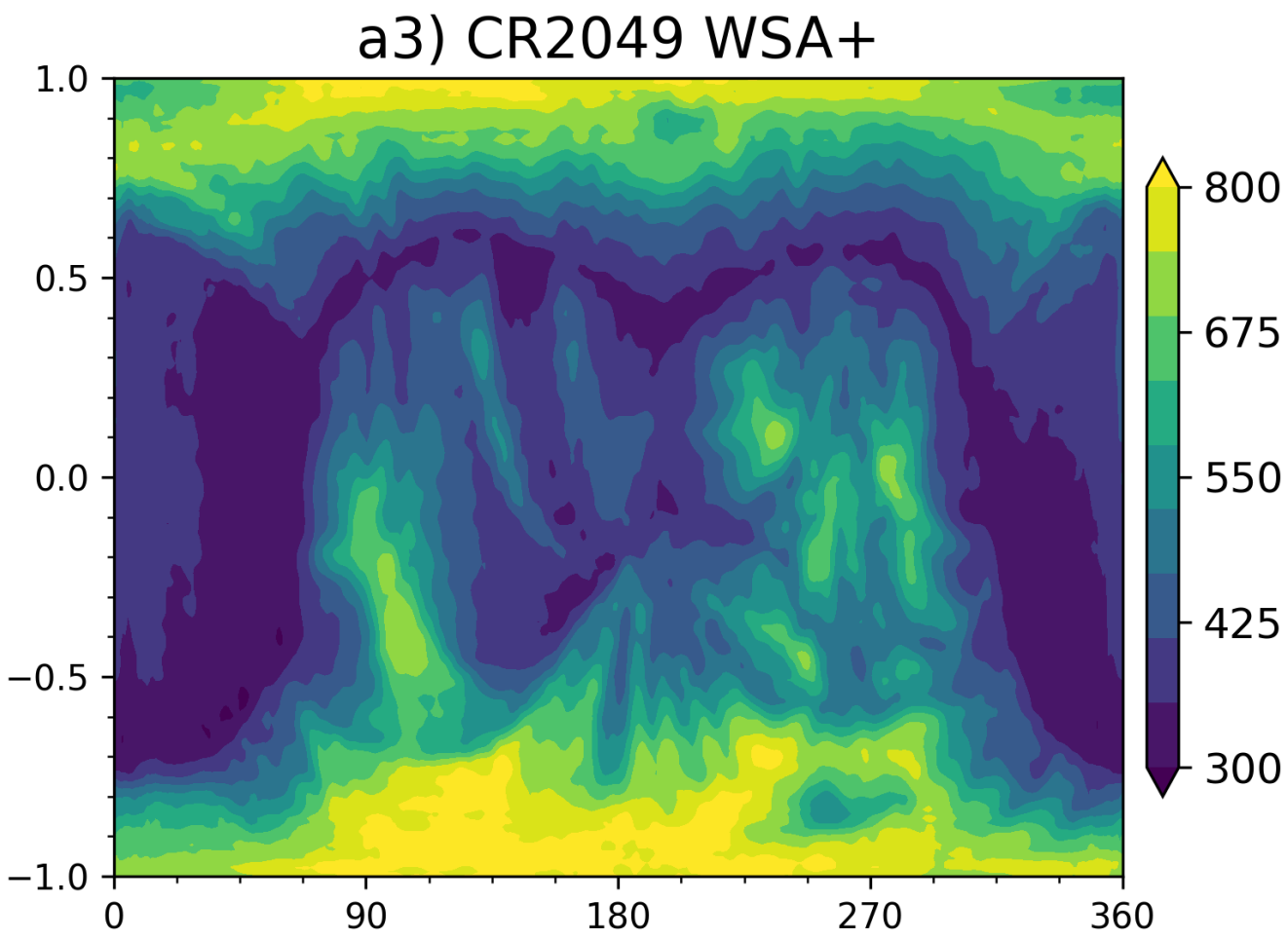
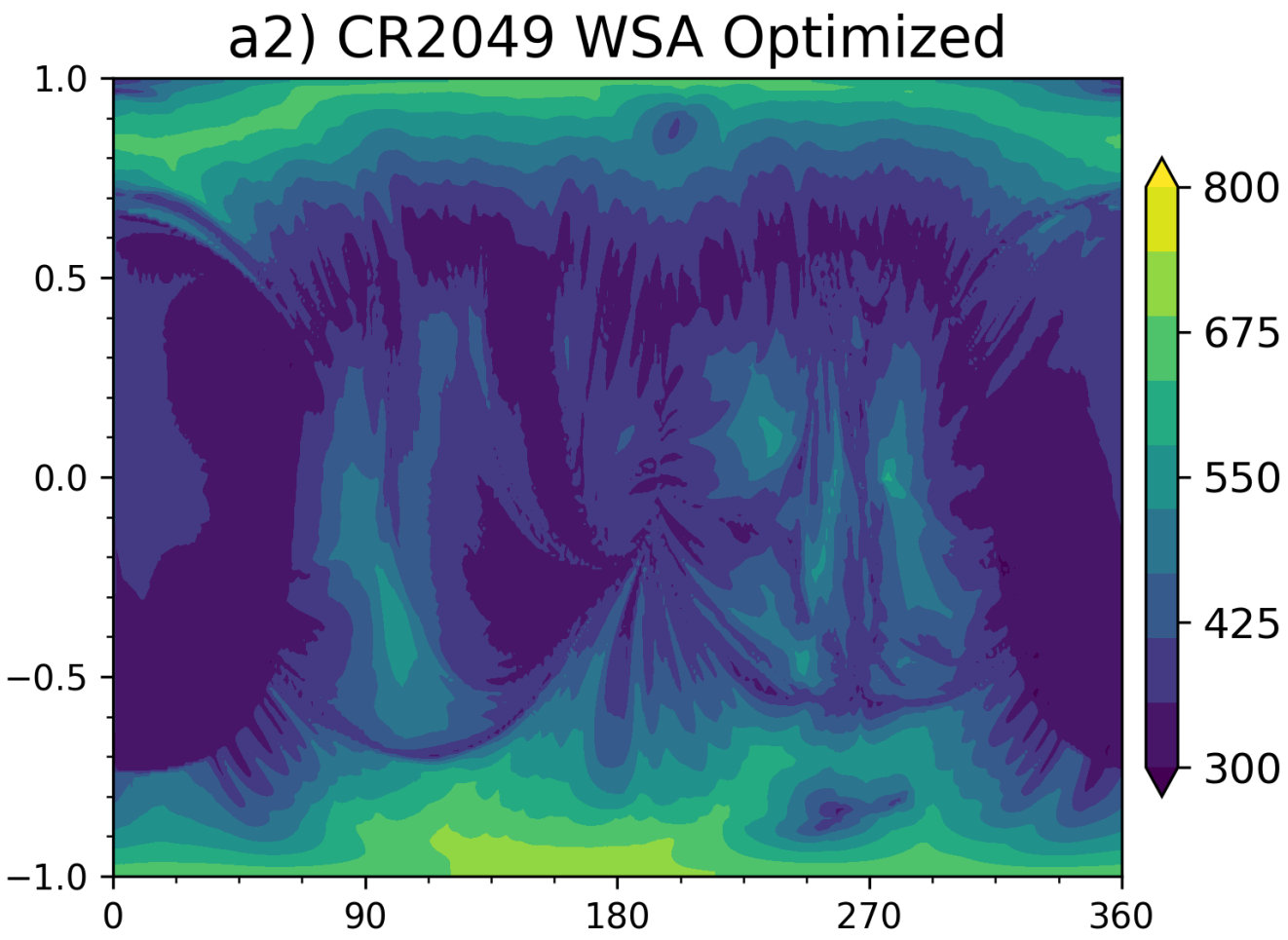
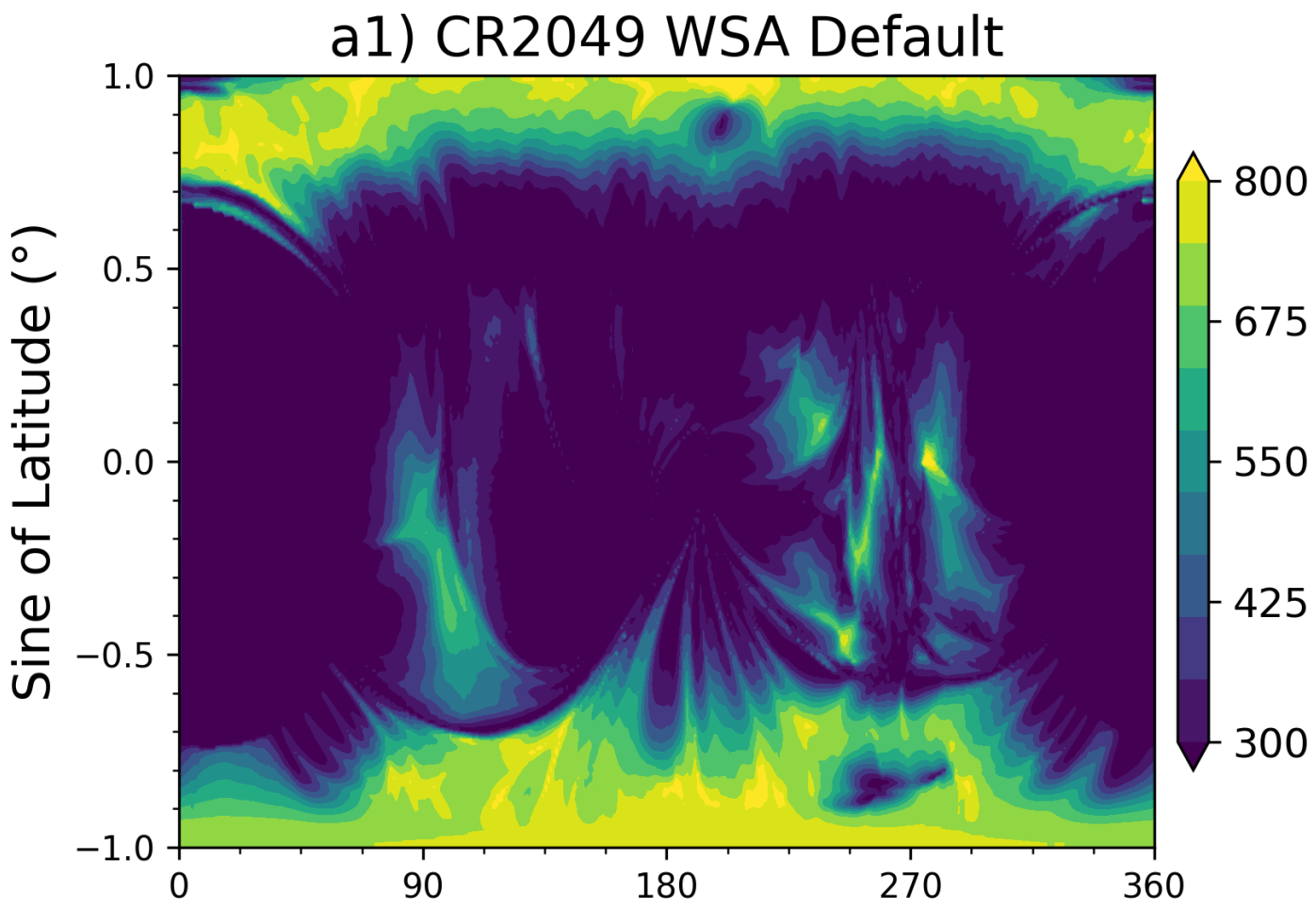
$$\alpha = 0.222, w = 0.028, \beta = 1.25,$$

$$a1 = 0.8, a2 = 3.0$$

Development of WSA+



Comparison with WSA

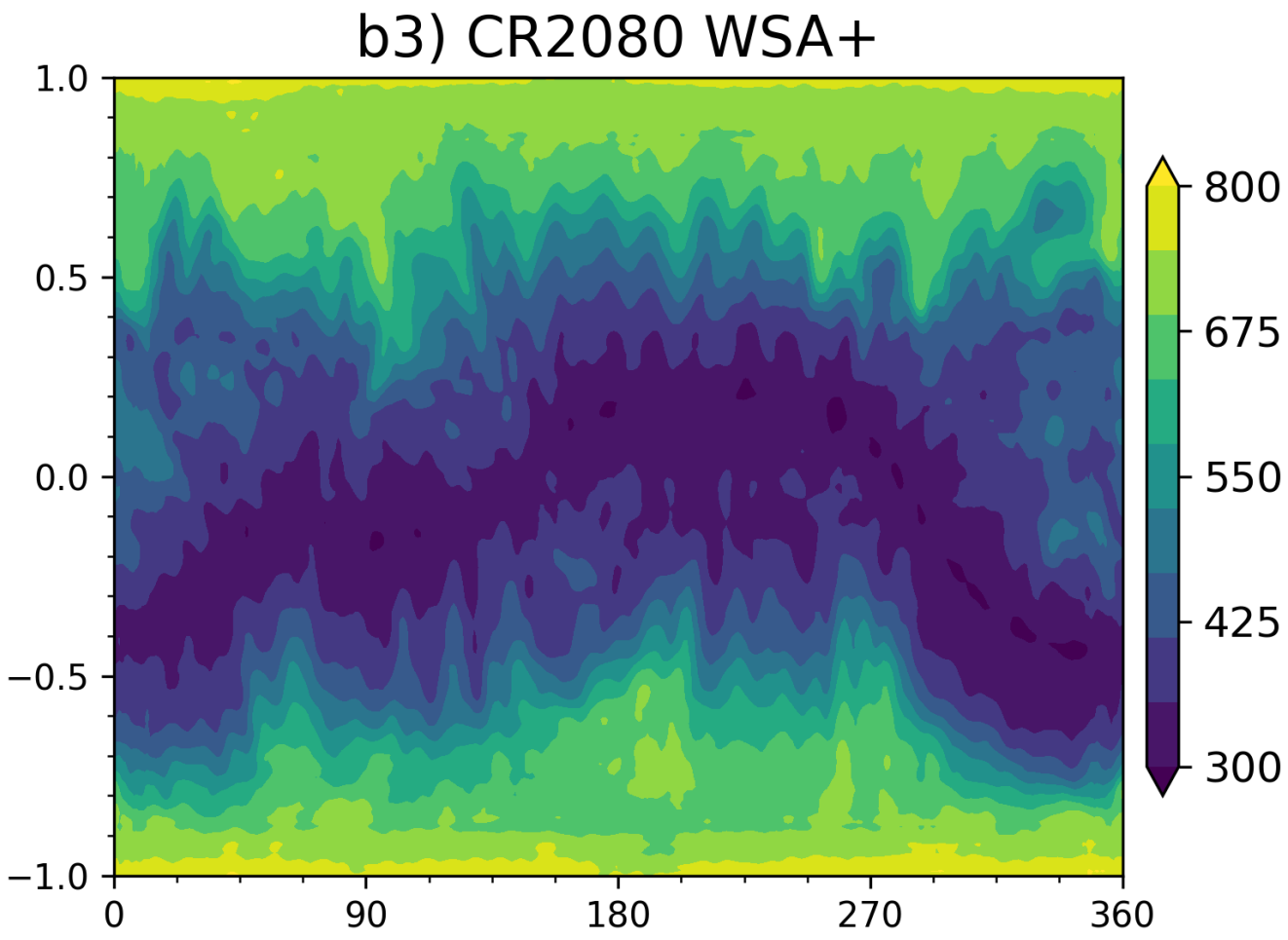
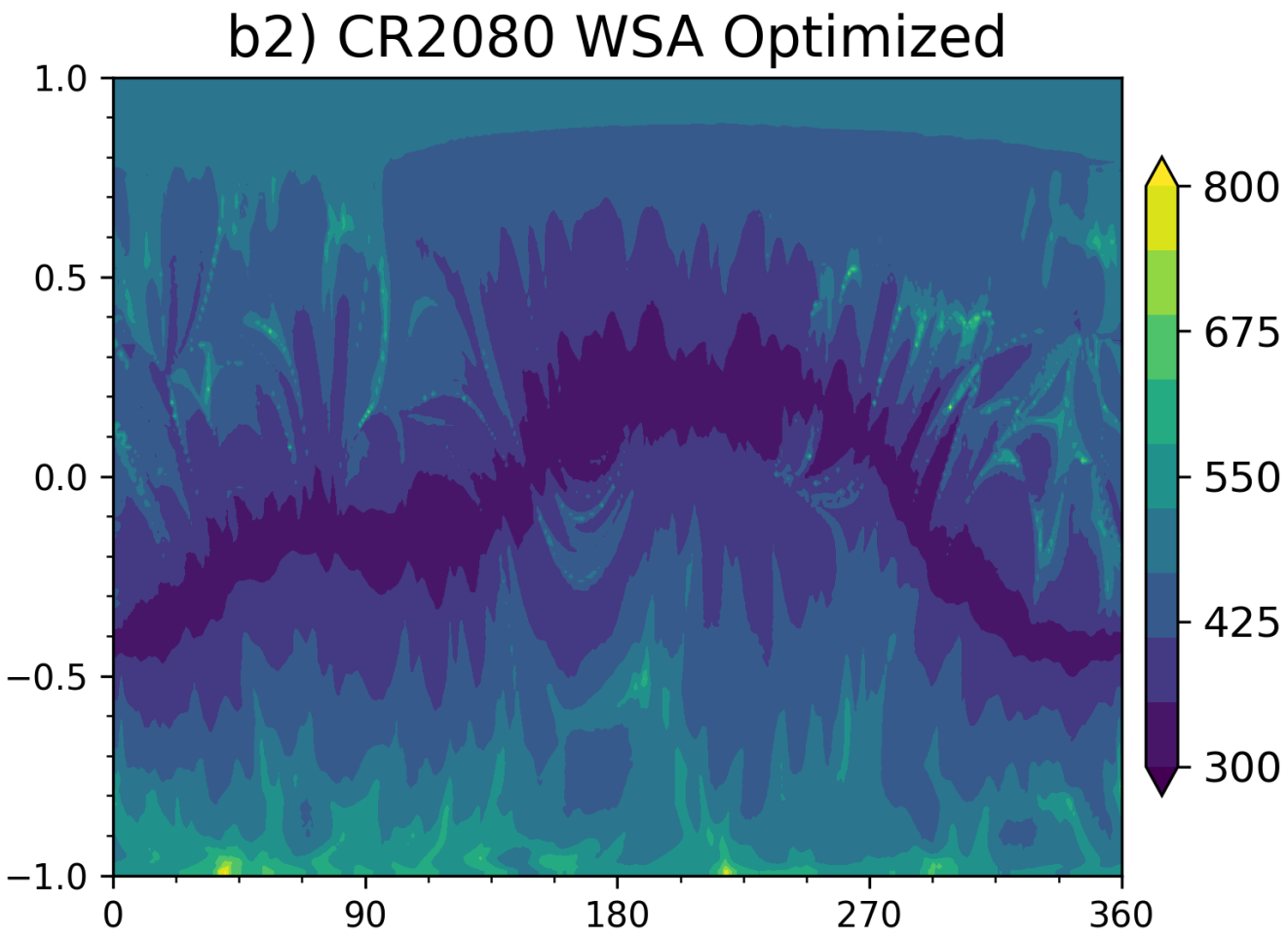
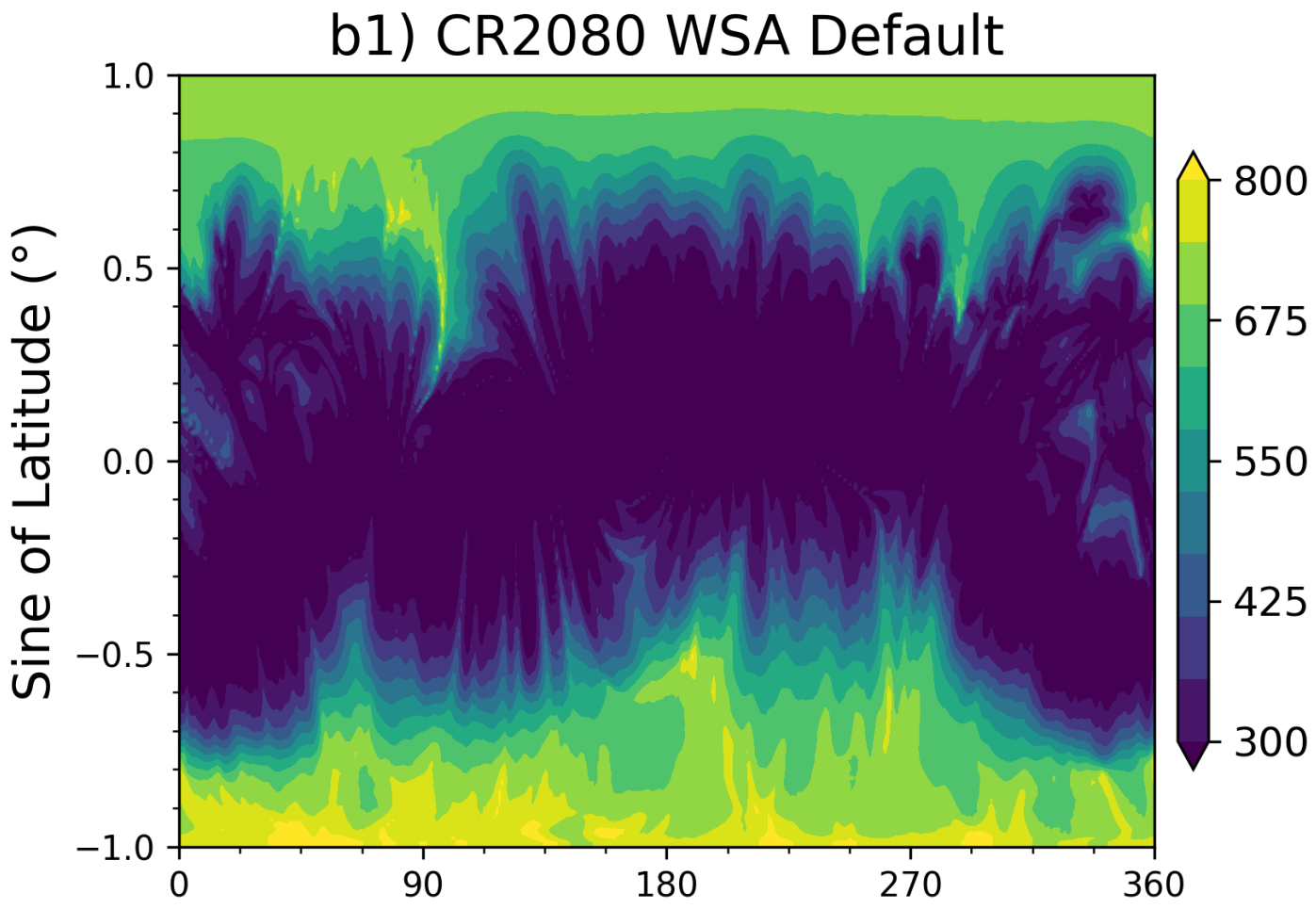


a4) Statistical Summary

CR2049 Analysis
=====

RMSE Comparison:
vs Default: 120.7 km/s
vs Optimized: 100.6 km/s

Correlation Comparision:
vs Default: 0.945
vs Optimized: 0.953



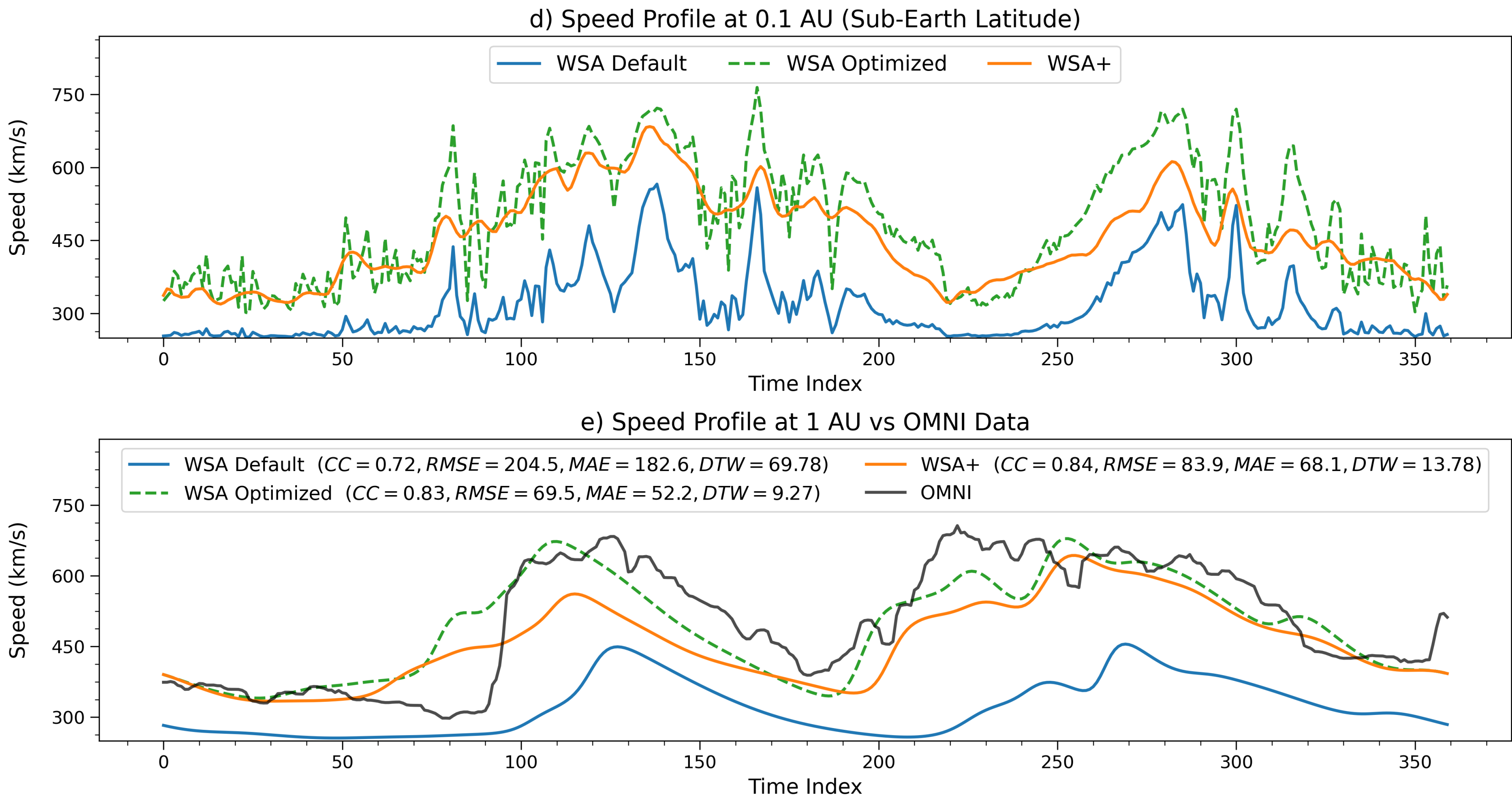
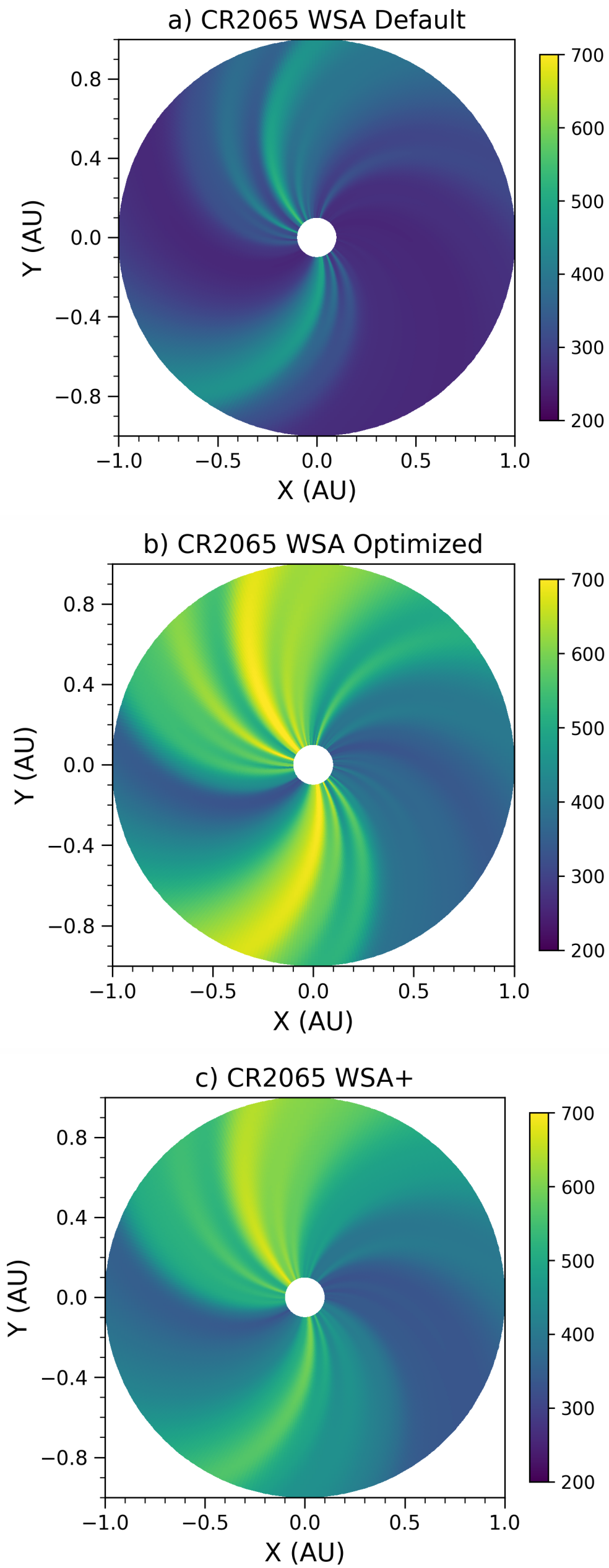
b4) Statistical Summary

CR2080 Analysis
=====

RMSE Comparison:
vs Default: 79.1 km/s
vs Optimized: 133.5 km/s

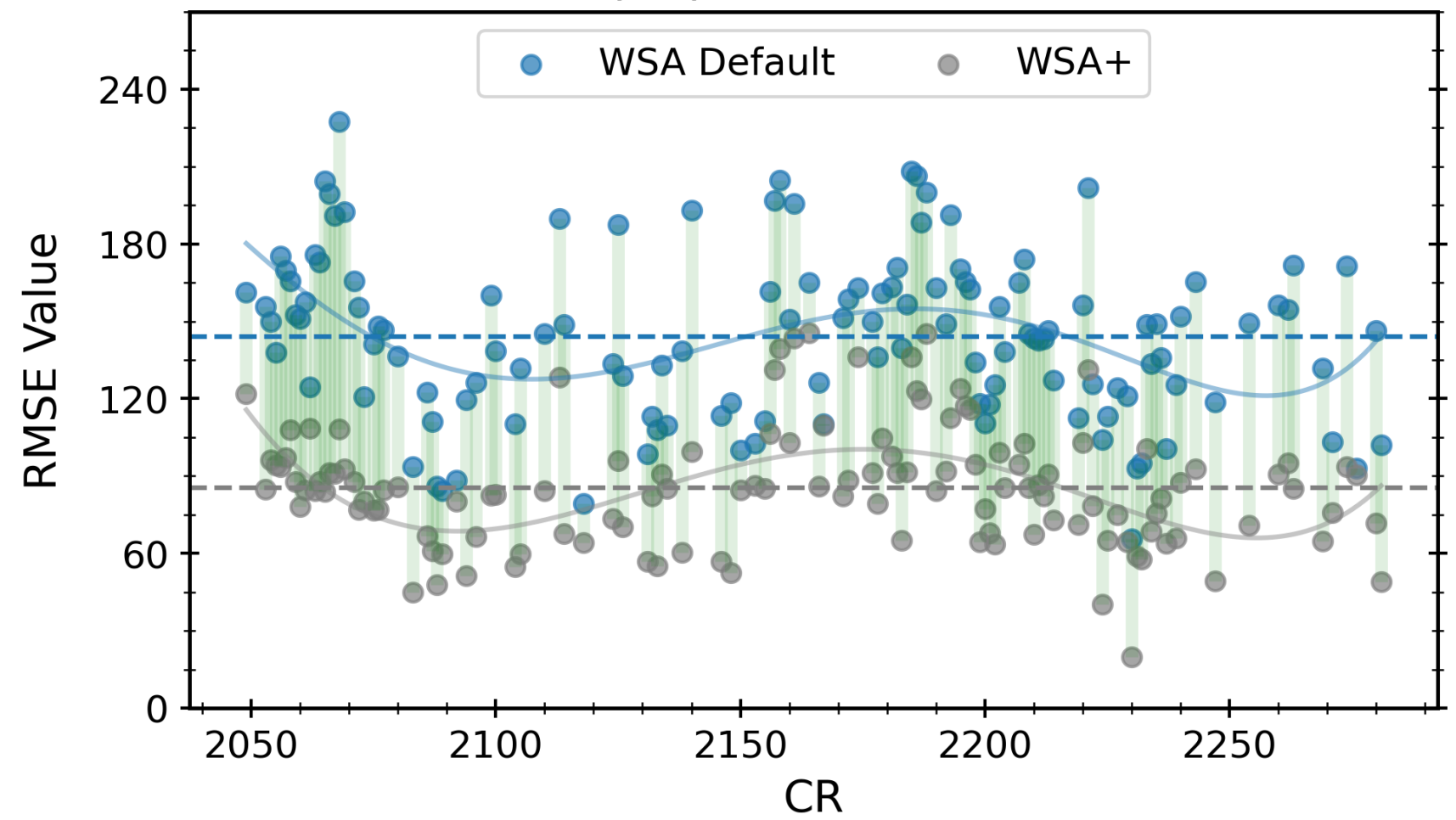
Correlation Comparision:
vs Default: 0.959
vs Optimized: 0.835

Comparison with OMNI

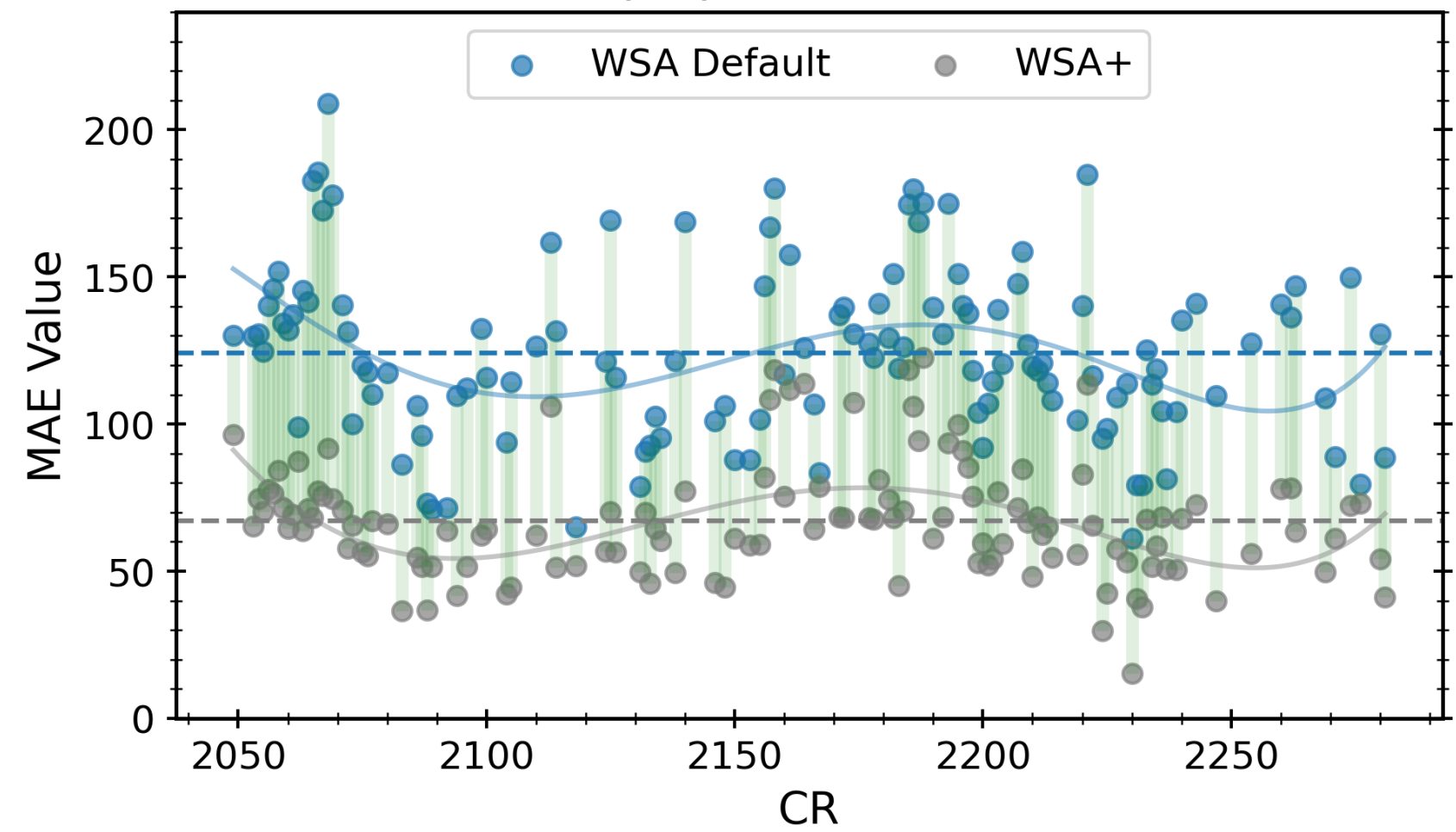


Long Term Analysis

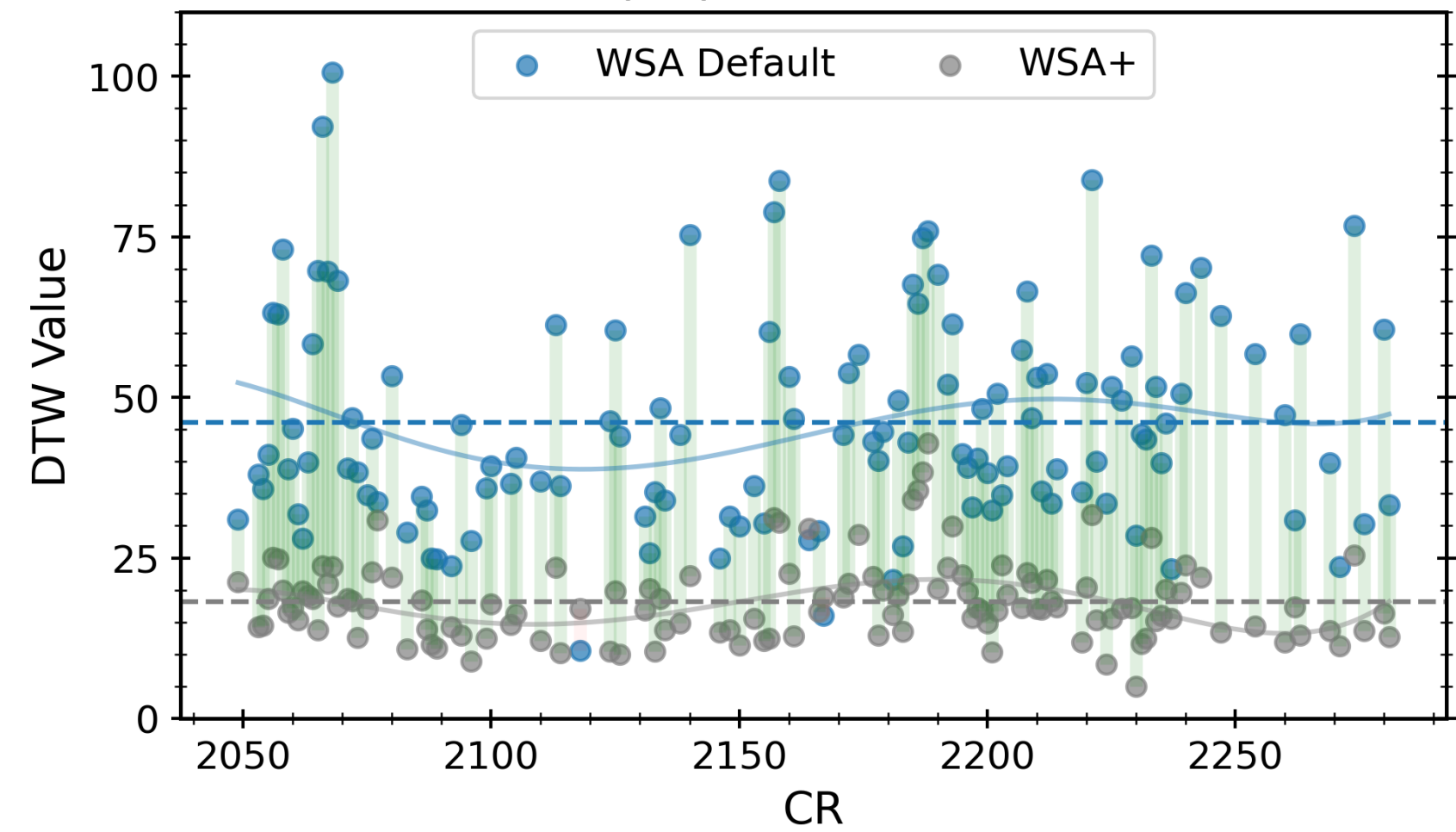
(a1) RMSE vs CR



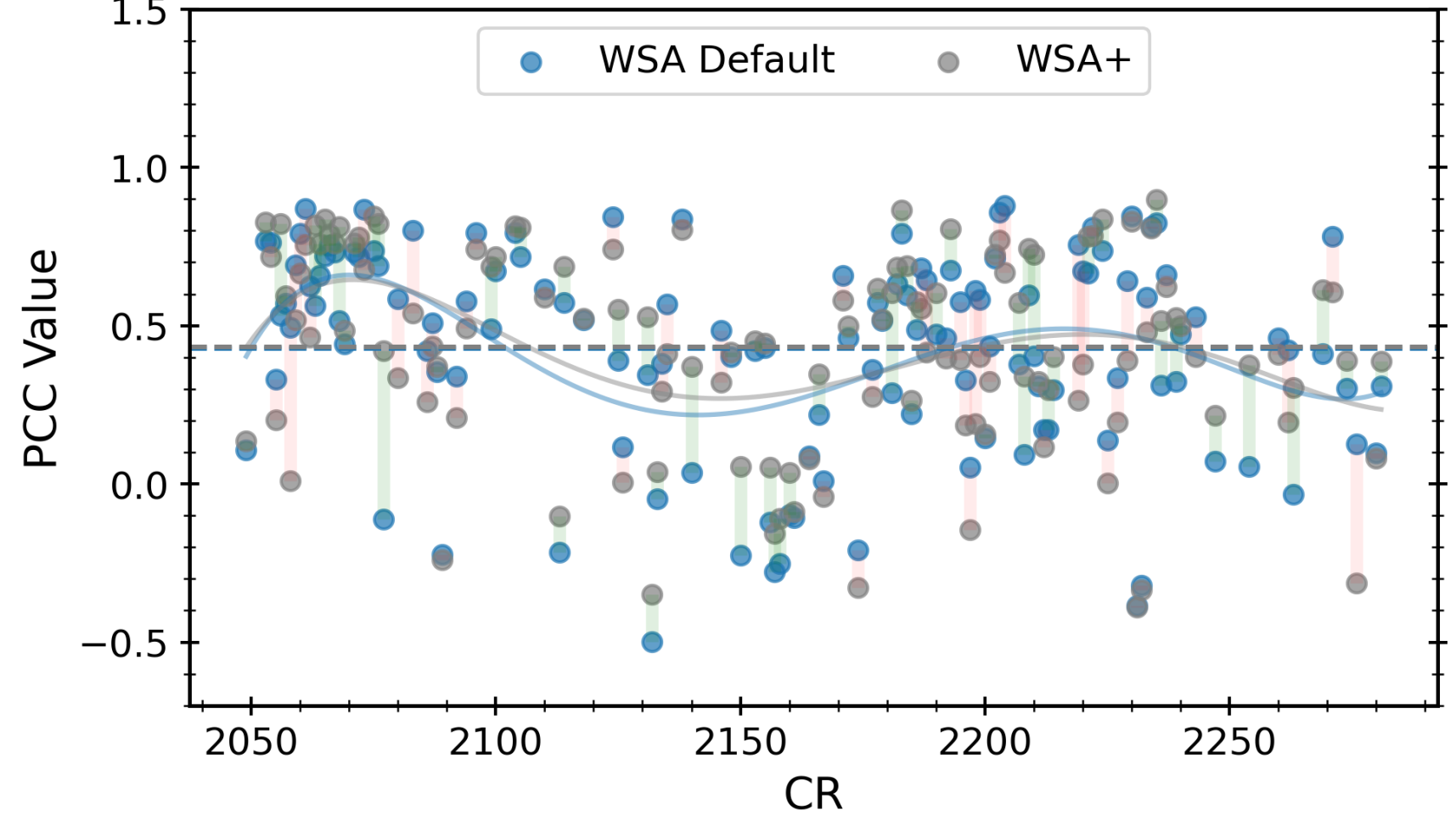
(a2) MAE vs CR



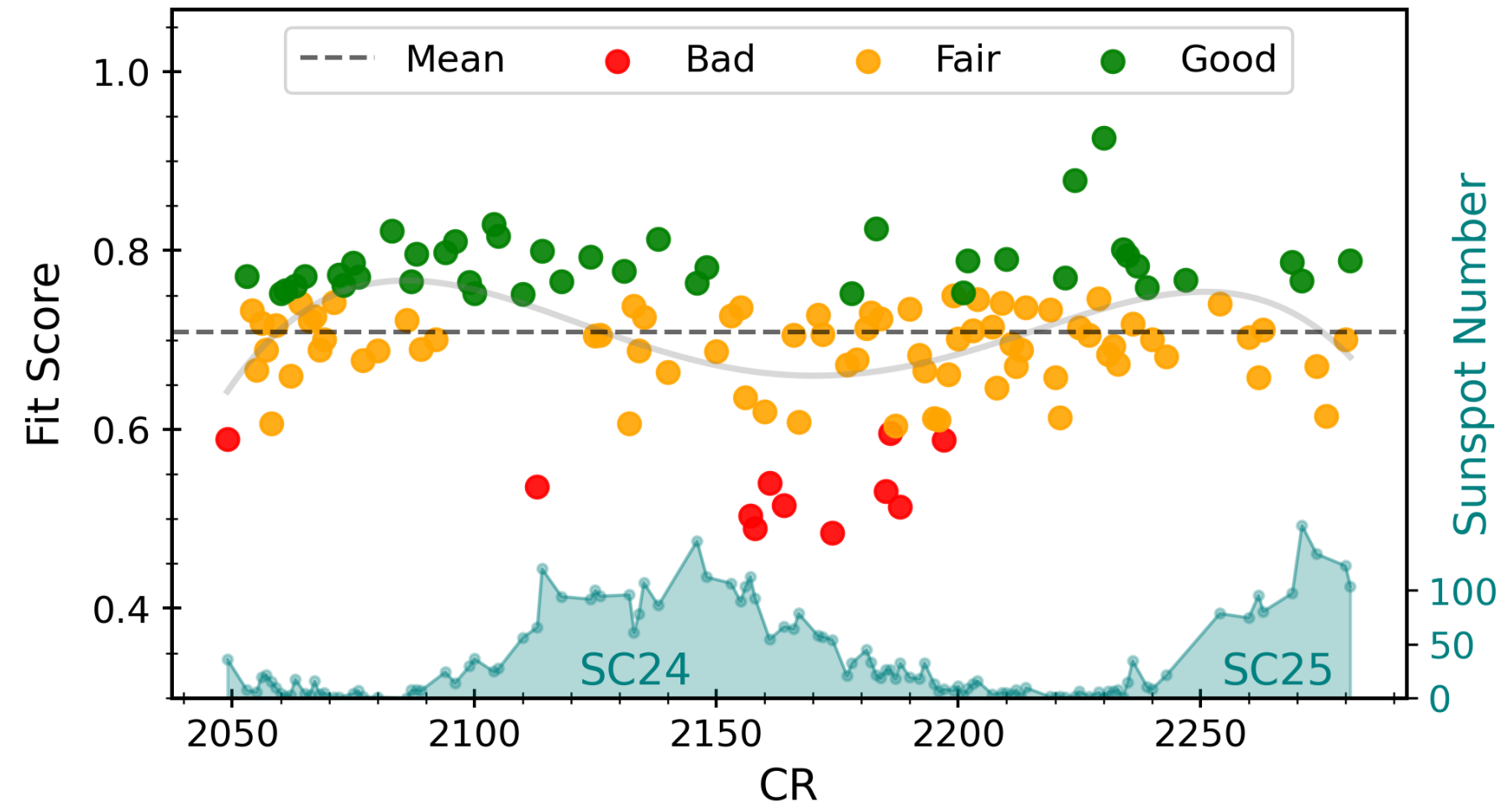
(a3) DTW vs CR



(a4) PCC vs CR



(b) Fit Score vs CR

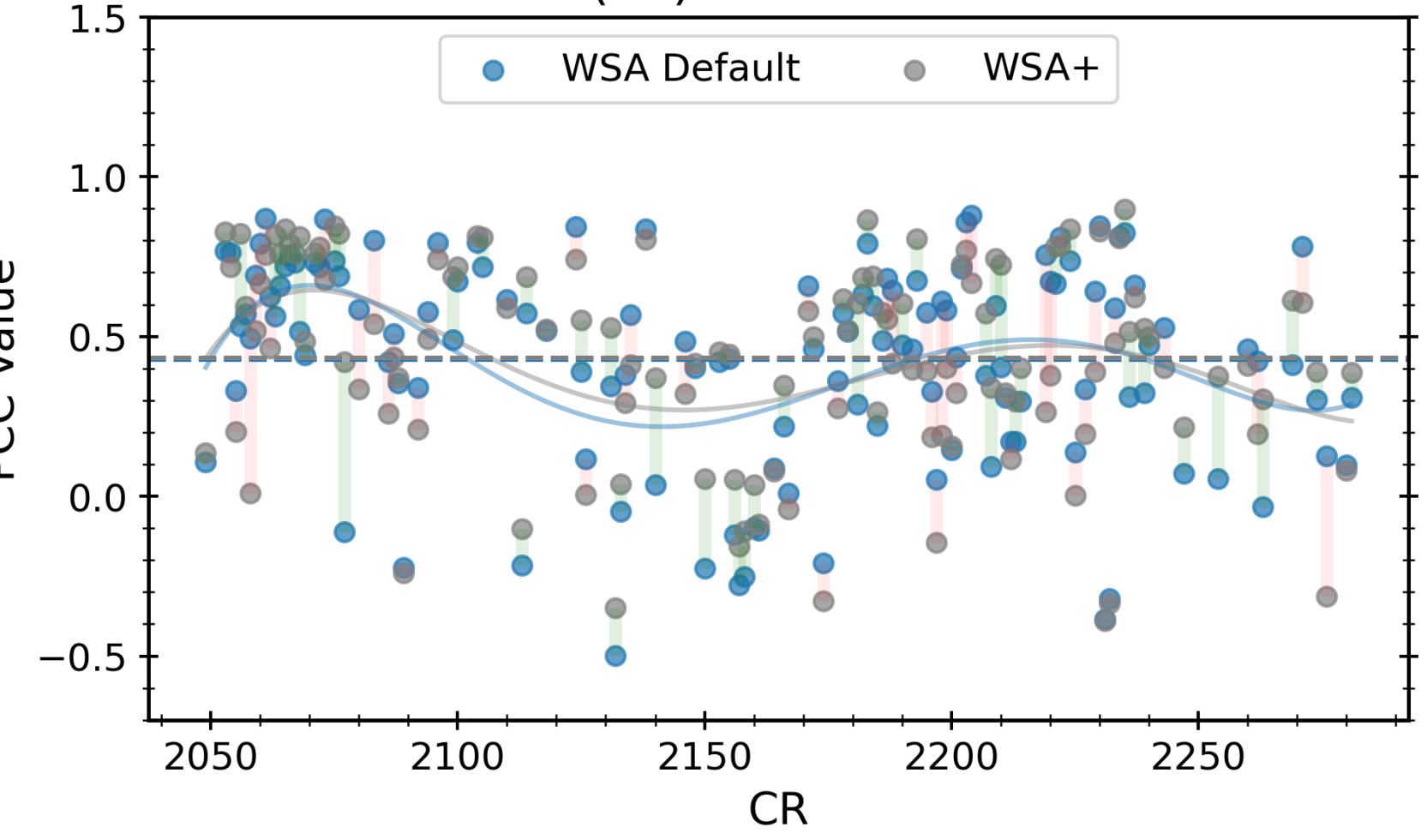


Performance Summary of 129 CRs

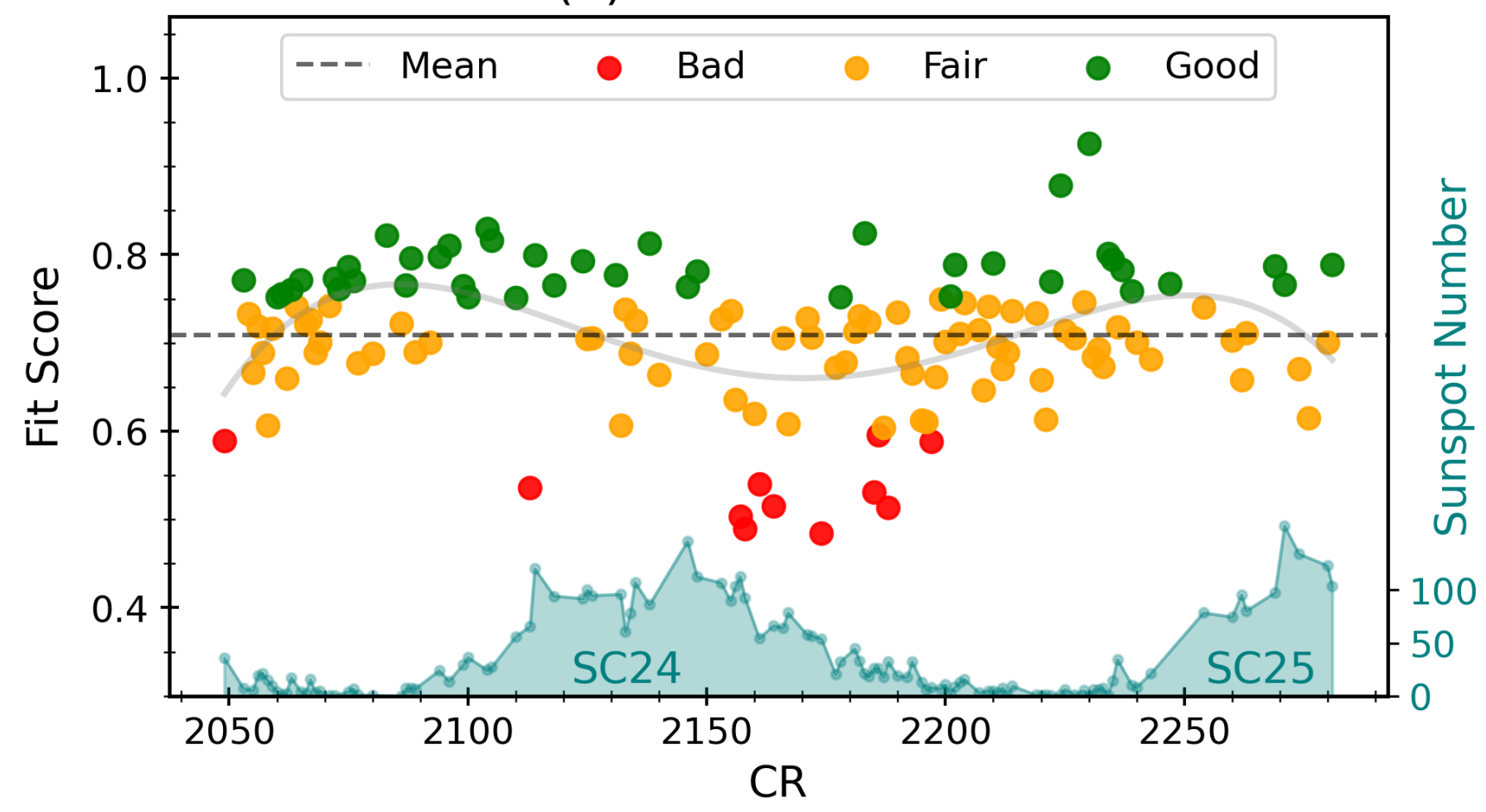
Avg. Metric	WSA	WSA+	Improvement
Fit Score	0.51	0.71	40.1%
DTW	46.07	18.21	60.5%
RMSE	144.00	85.53	40.6%
MAE	124.15	67.10	45.9%
PCC	0.43	0.43	1.3%

Long Term Analysis

(a4) PCC vs CR



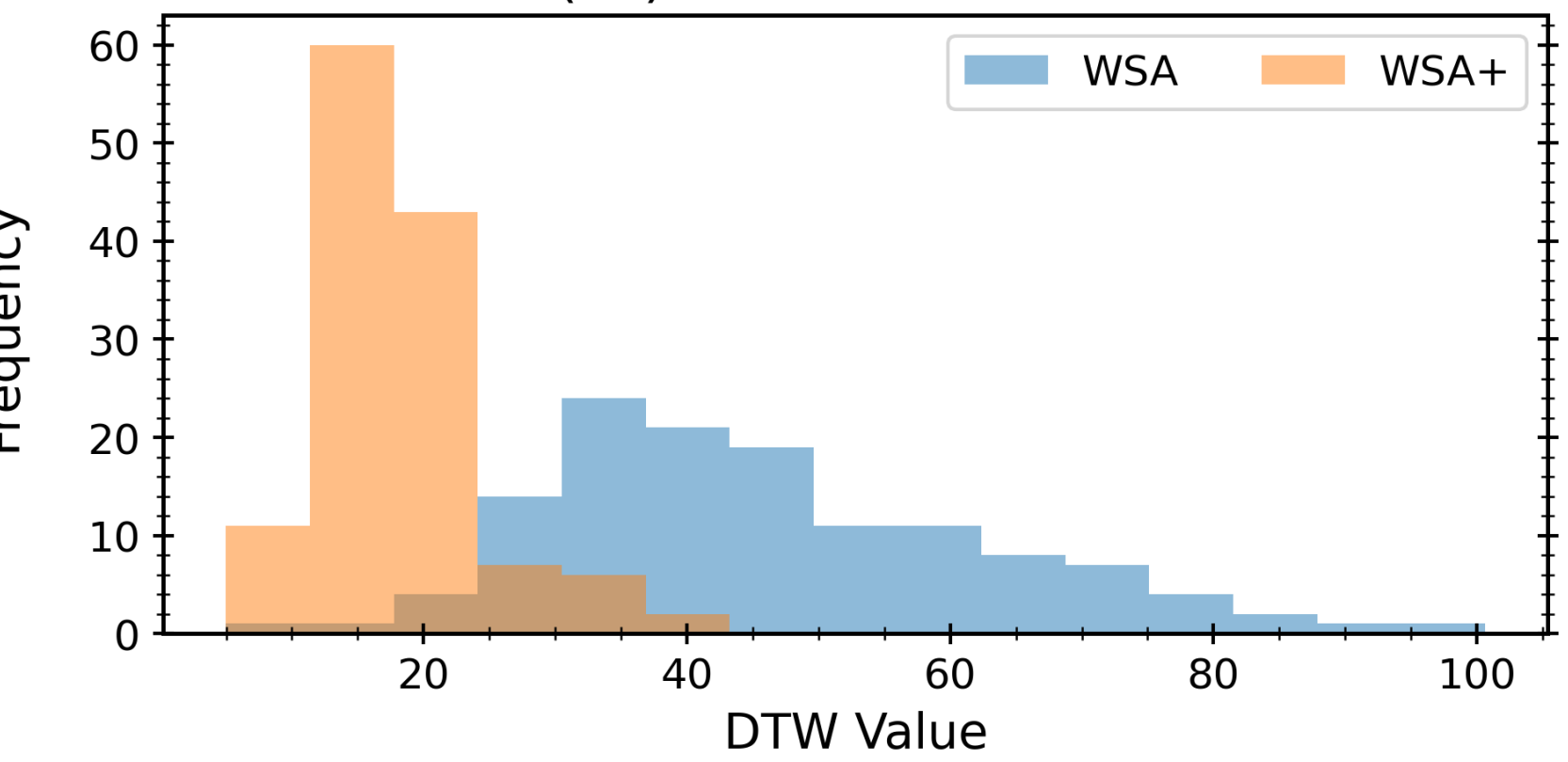
(b) Fit Score vs CR



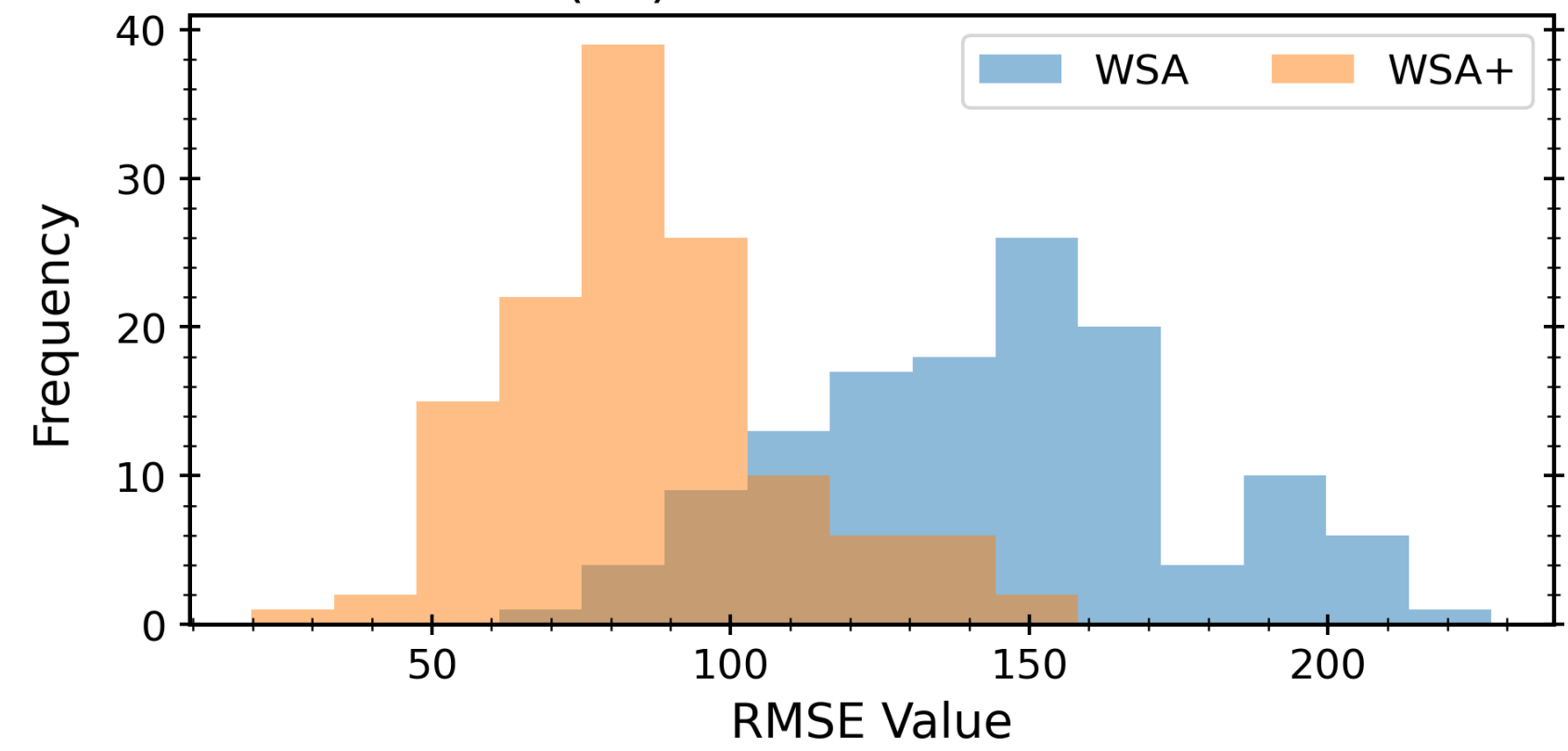
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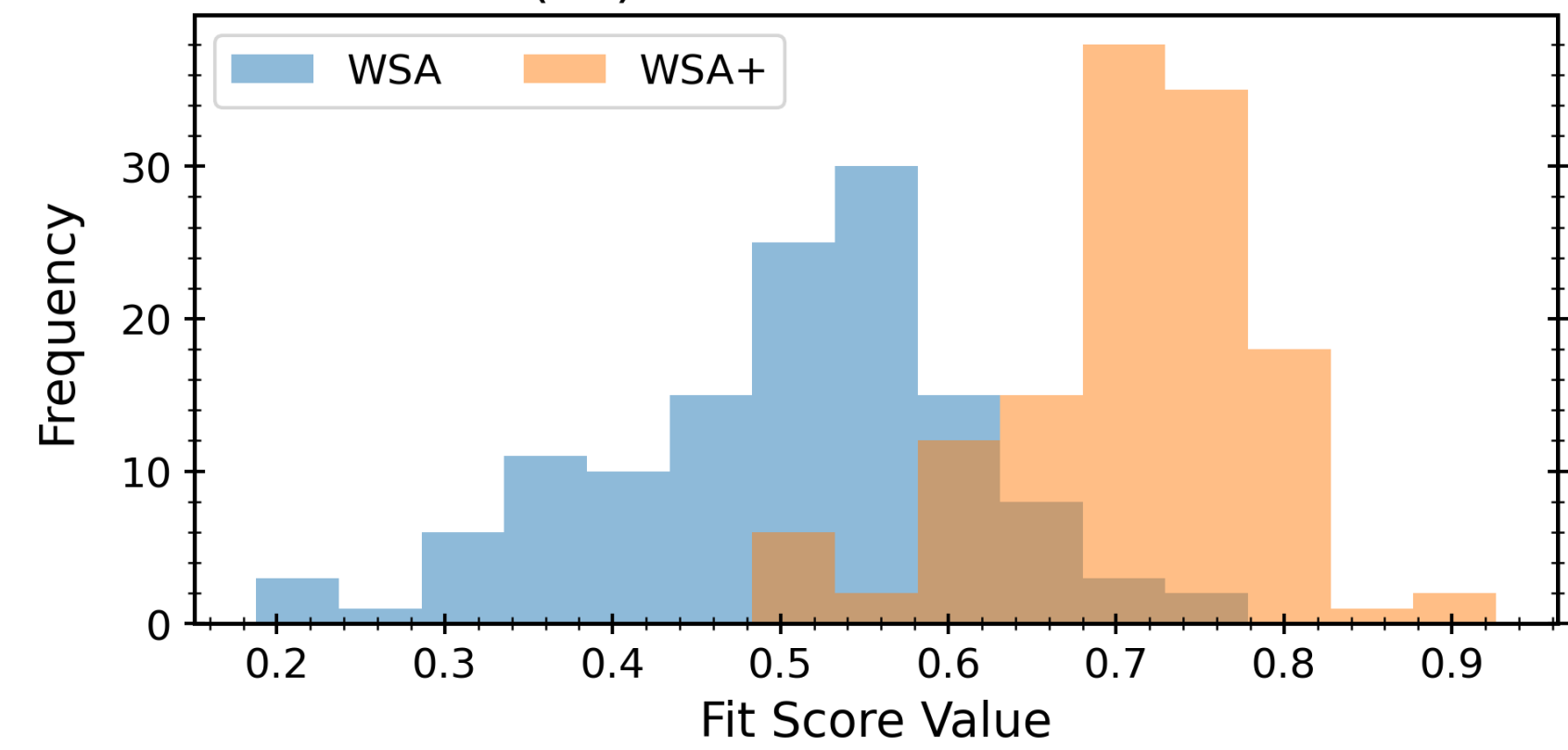
(c1) DTW Distribution



(c2) RMSE Distribution



(c3) Fit Score Distribution



Python Package



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wsaplus 0.1.1

`pip install wsaplus`

📄

✓

Latest version

Released: Aug 18, 2025

WSA+: generate solar wind speed maps from synoptic magnetograms

arXiv > astro-ph > arXiv:2509.06181

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[Submitted on 7 Sep 2025]

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[Prateek Mayank](#), [Enrico Camporeale](#), [Arpit K. Shrivastav](#), [Thomas E. Berger](#), [Charles N. Arge](#)

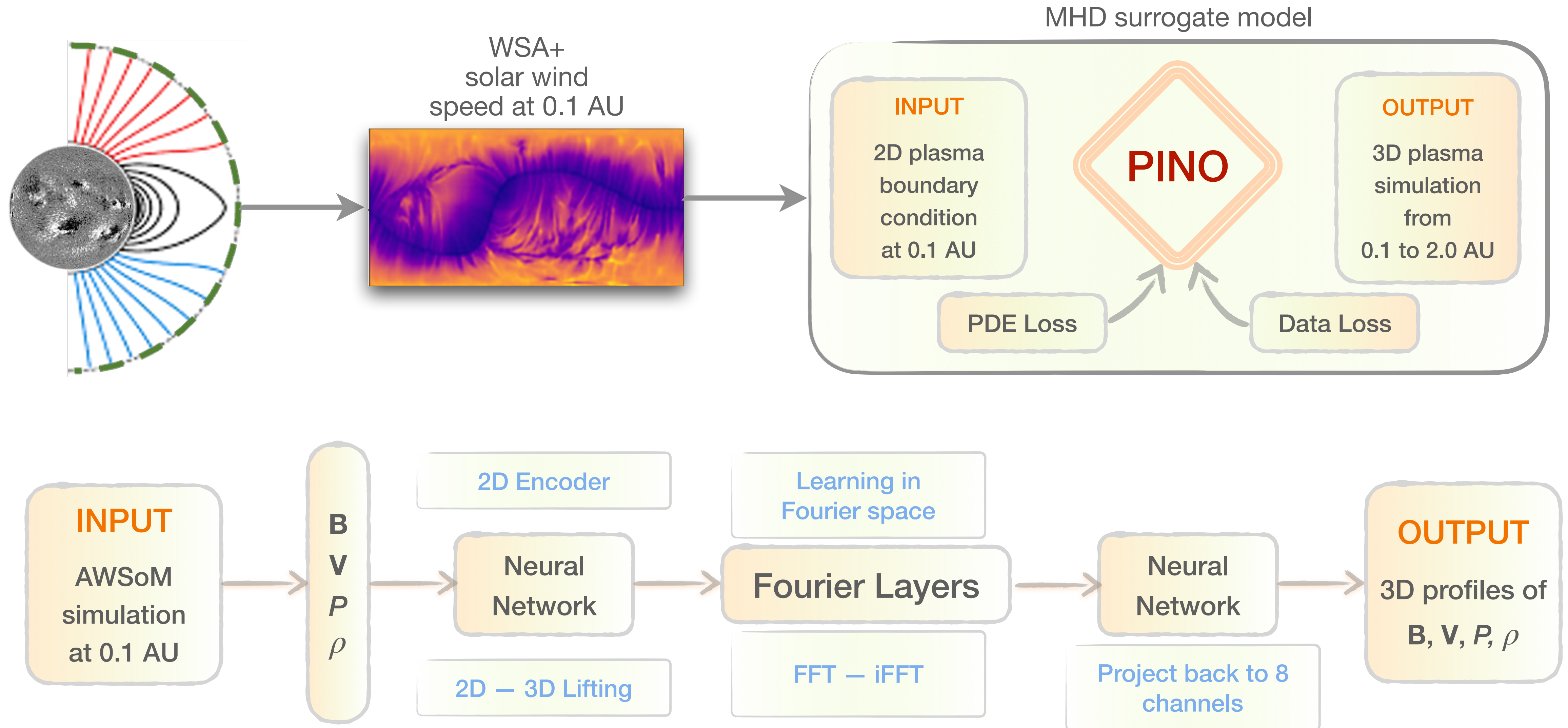
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MHD Surrogate Pipeline



Initial Results ...

ML

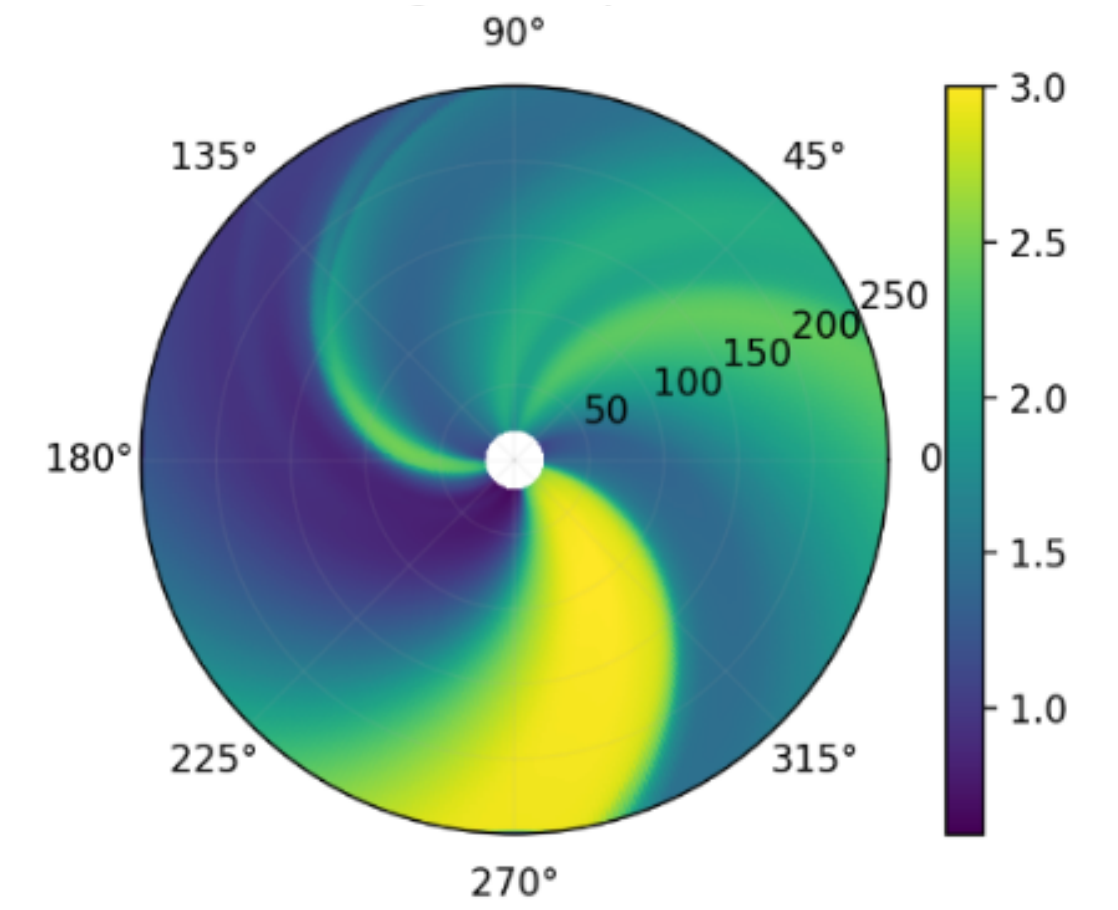
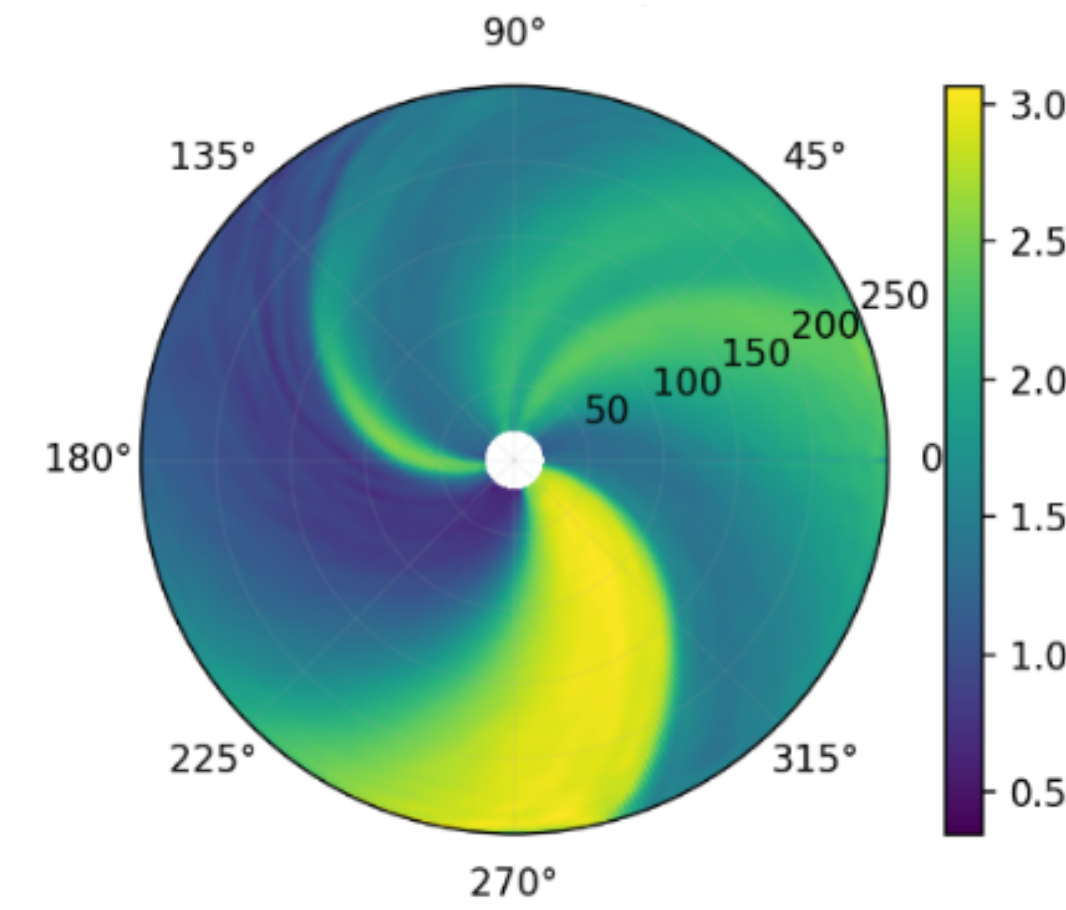
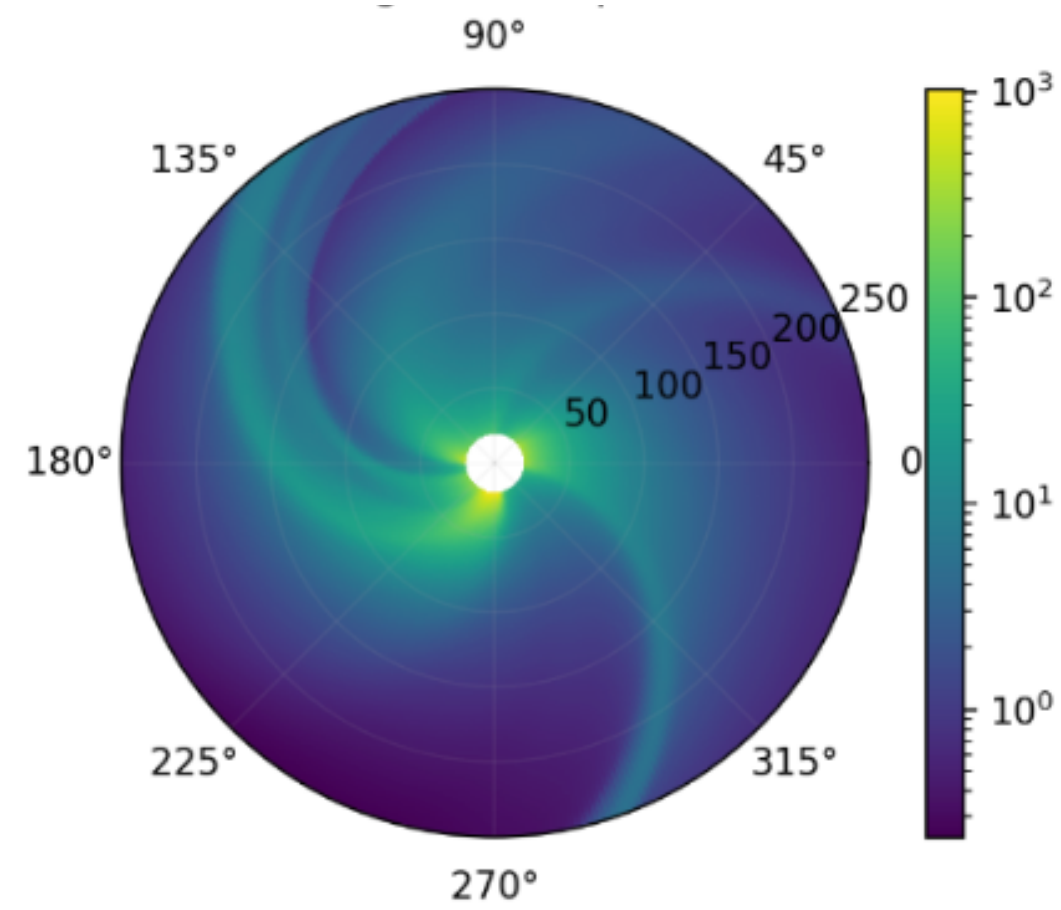
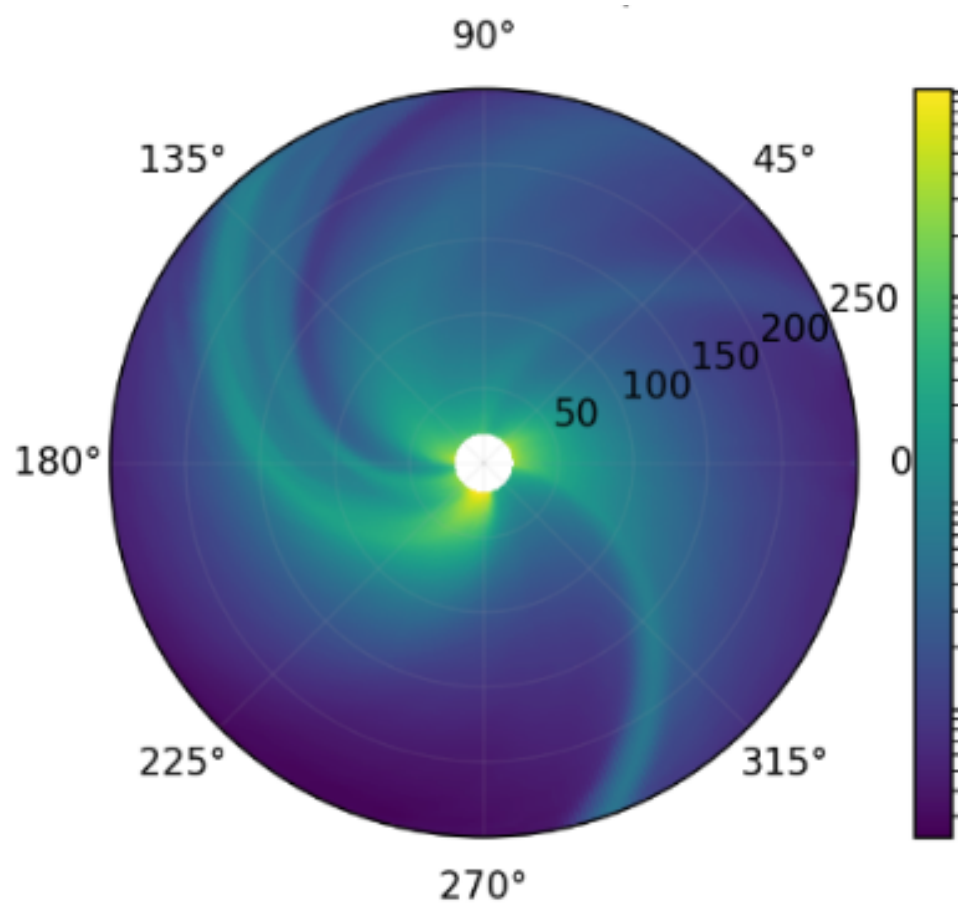
Density

MHD

ML

Speed

MHD



ML

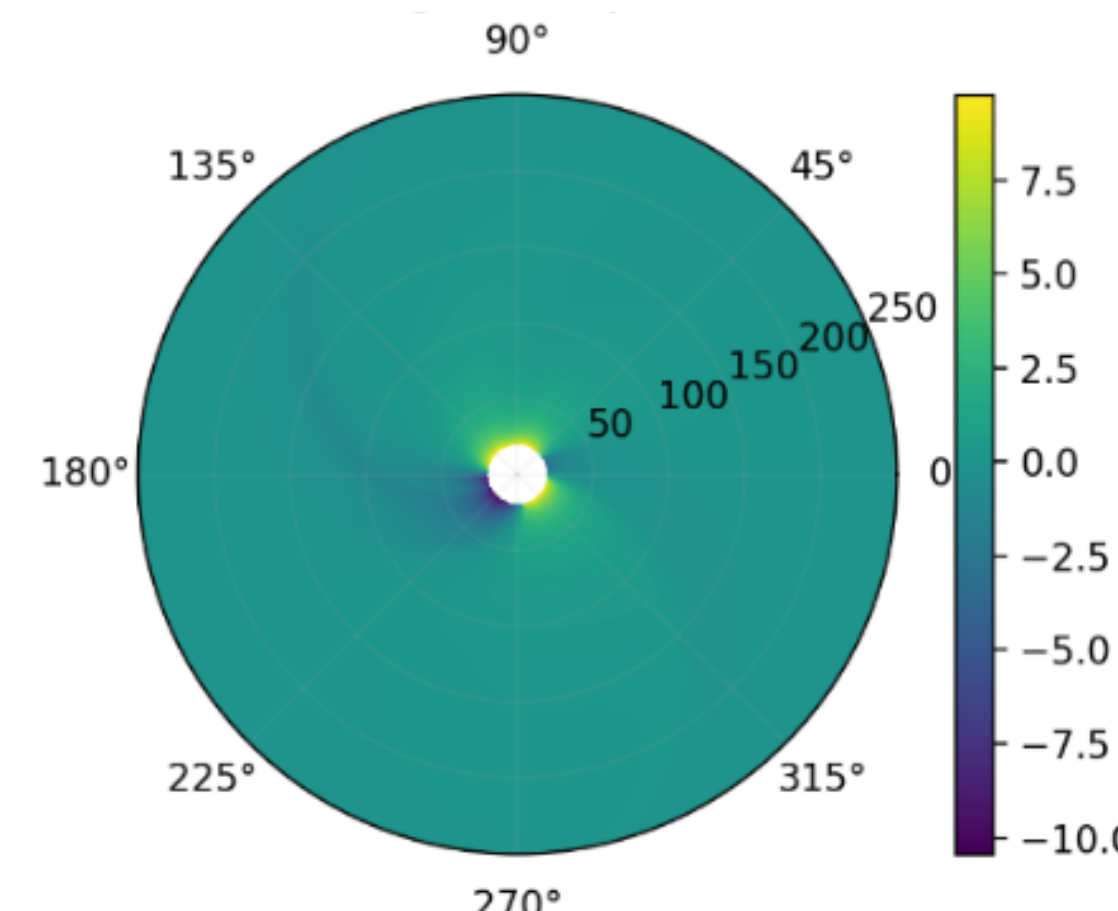
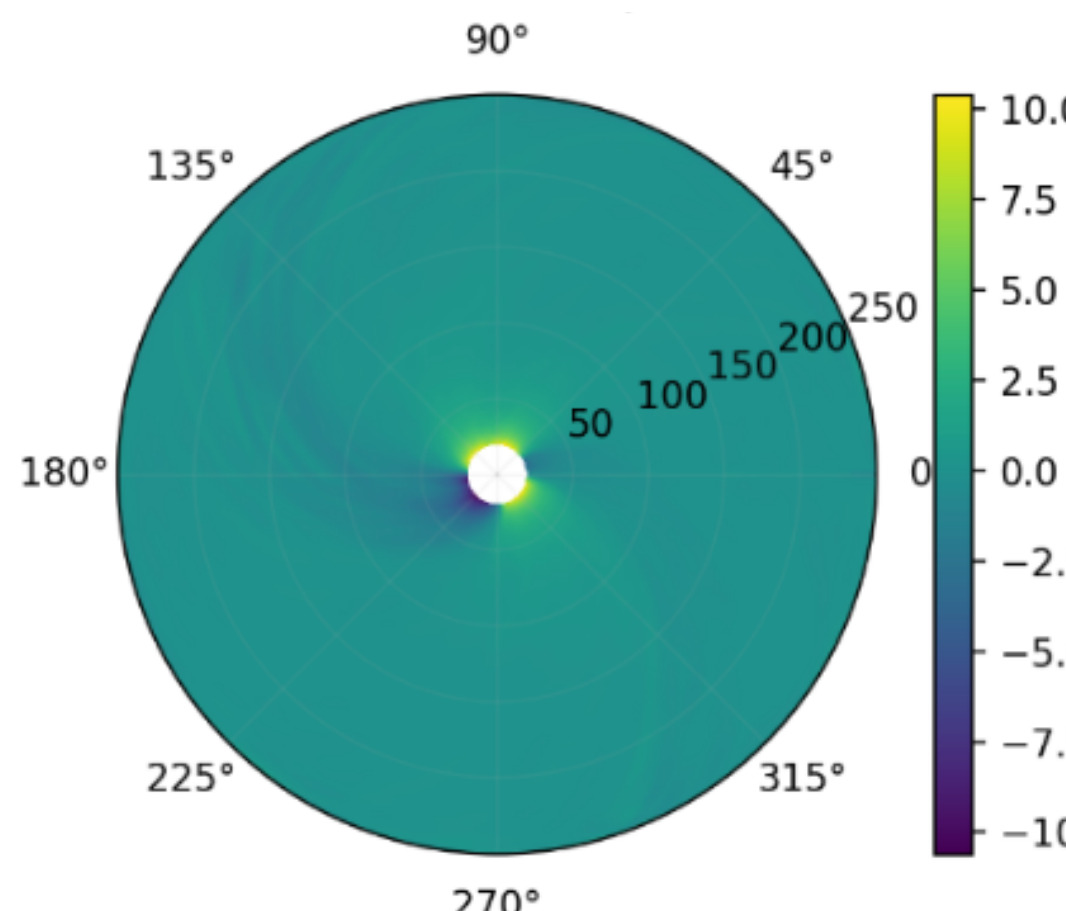
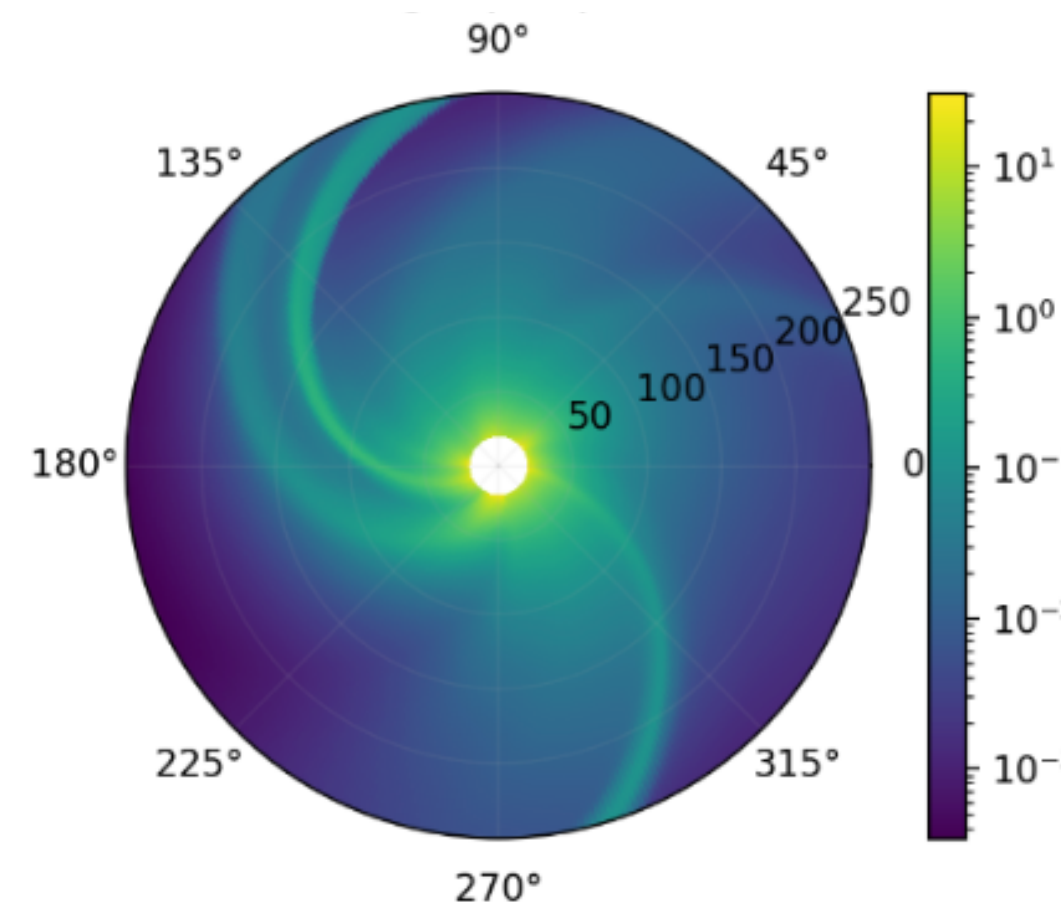
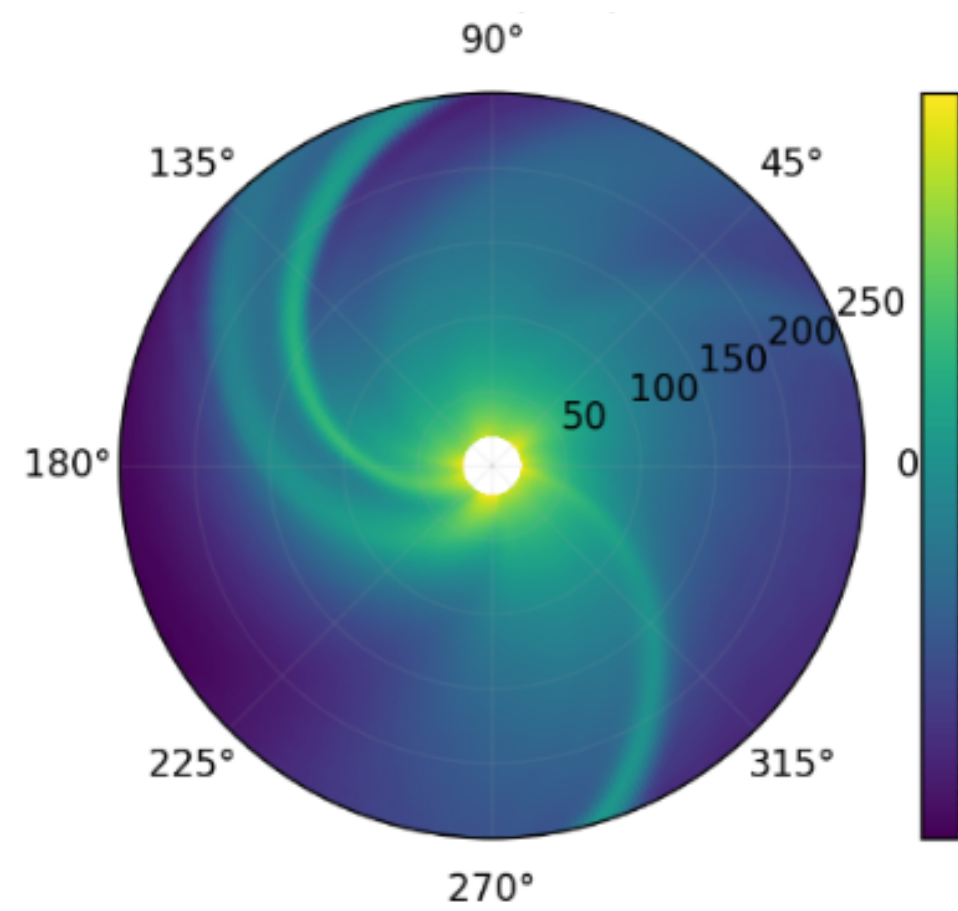
Pressure

MHD

ML

Magnetic Field

MHD



Initial Results ...

ML

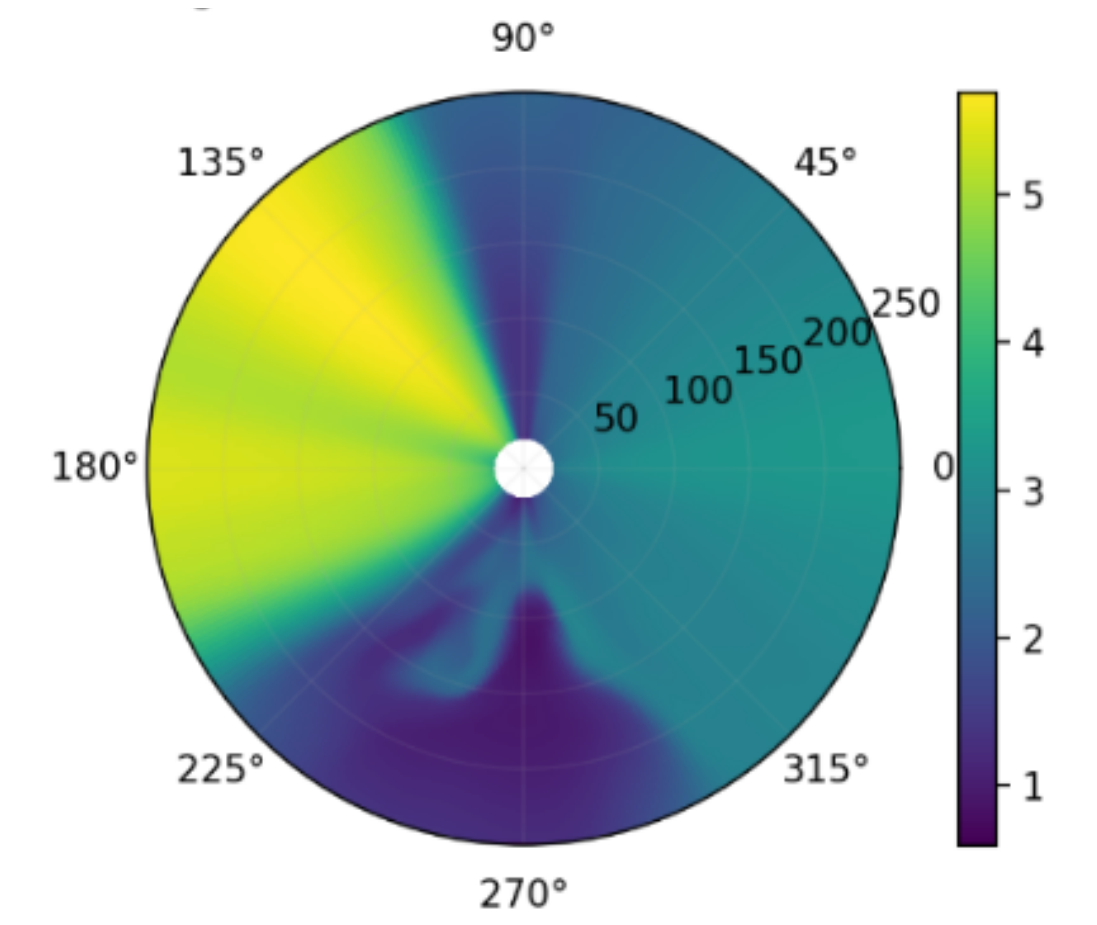
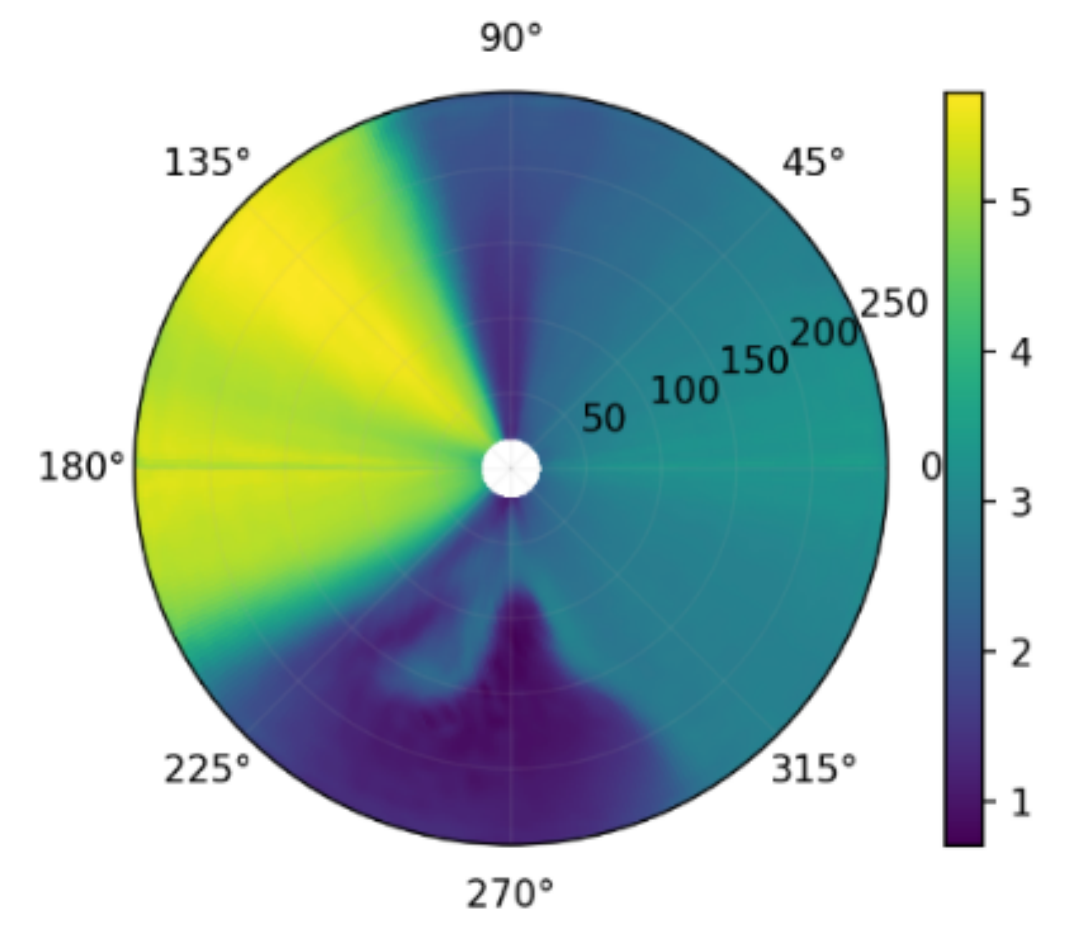
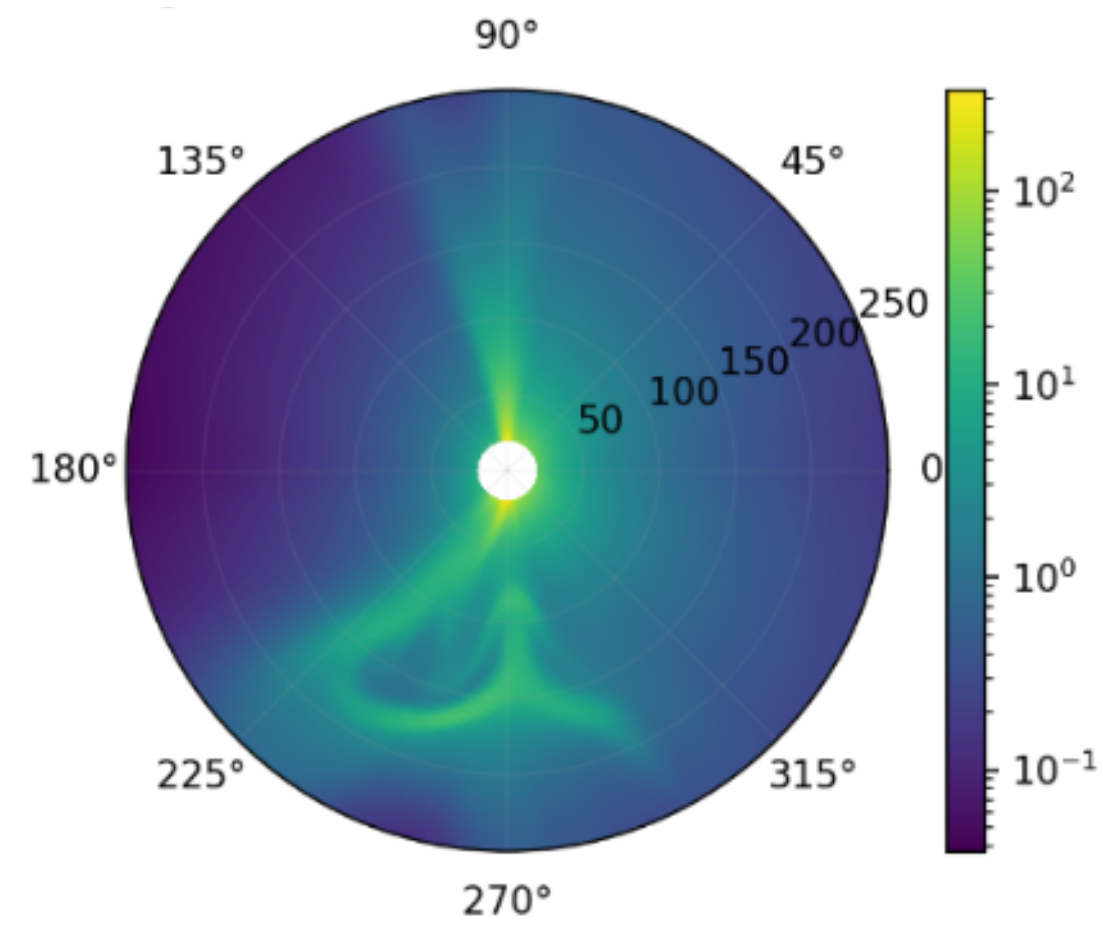
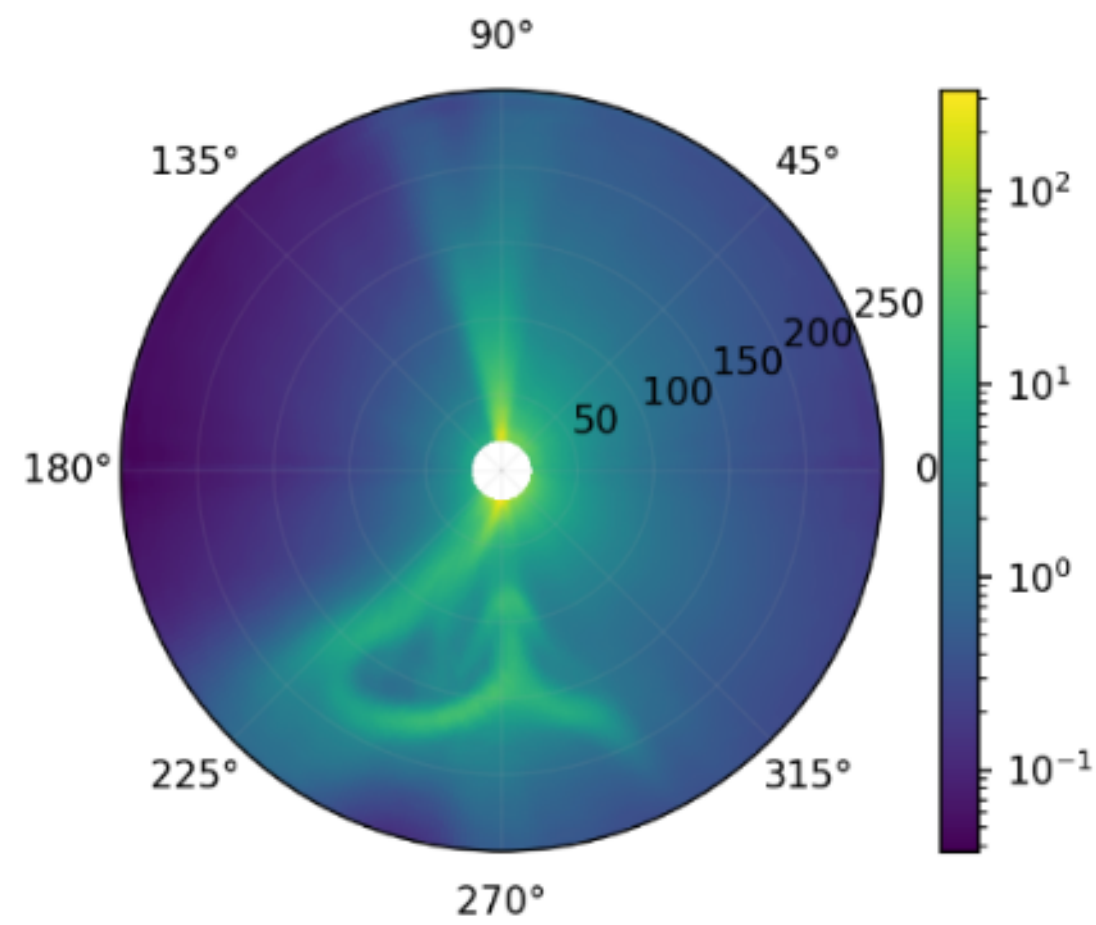
Density

MHD

ML

Speed

MHD



ML

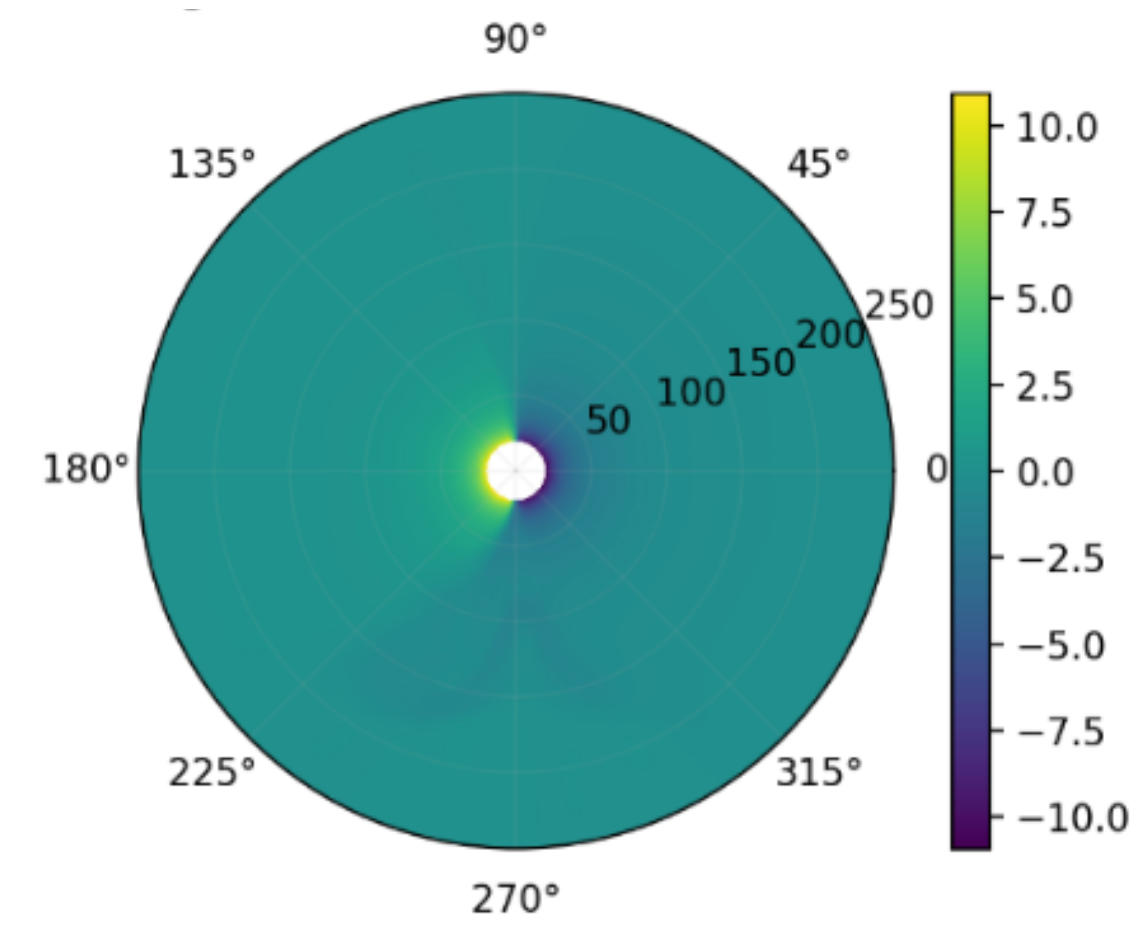
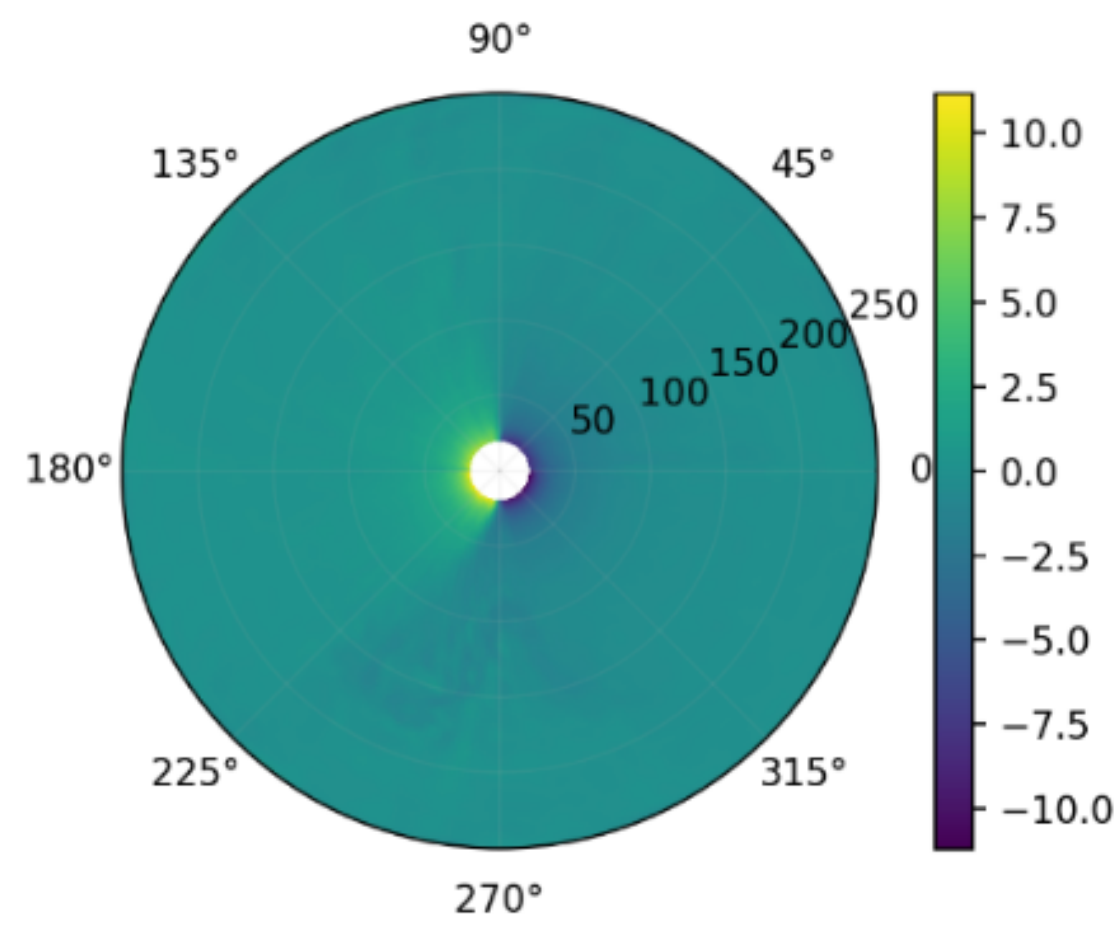
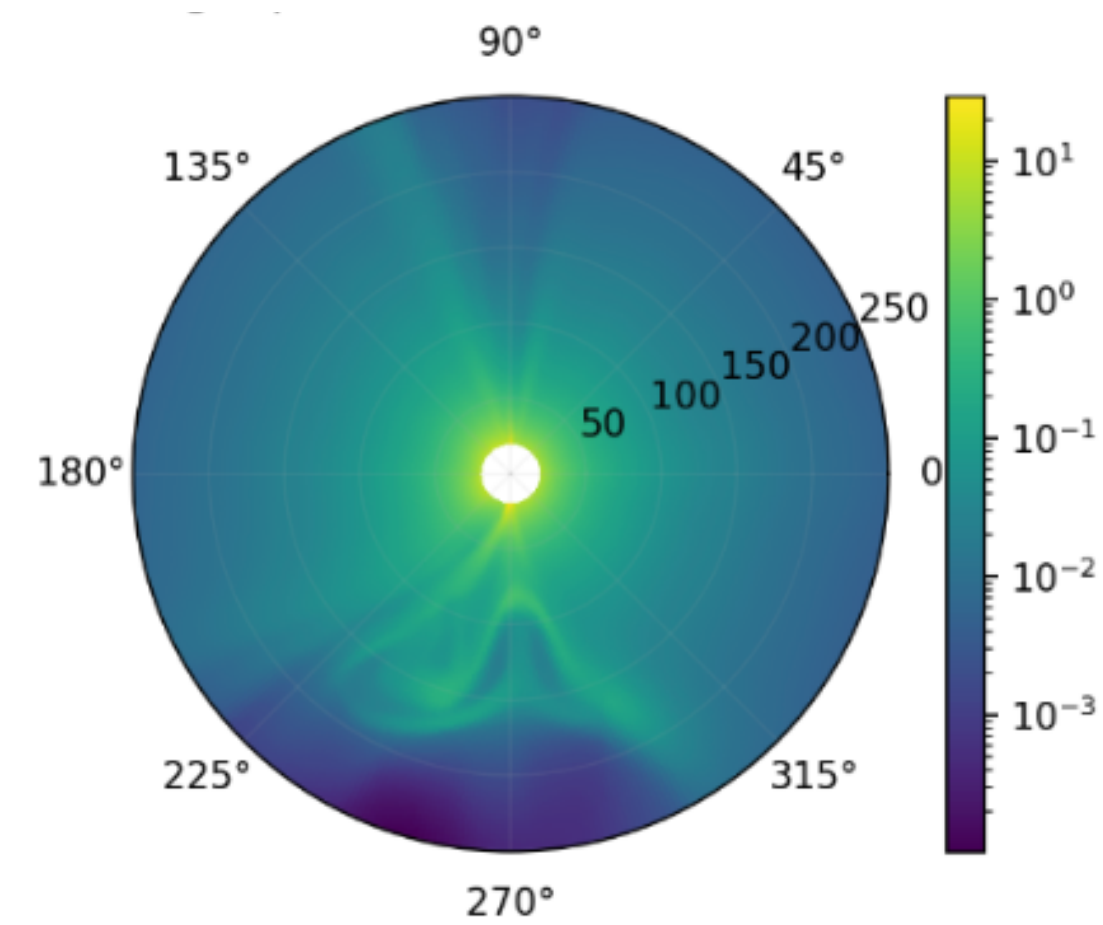
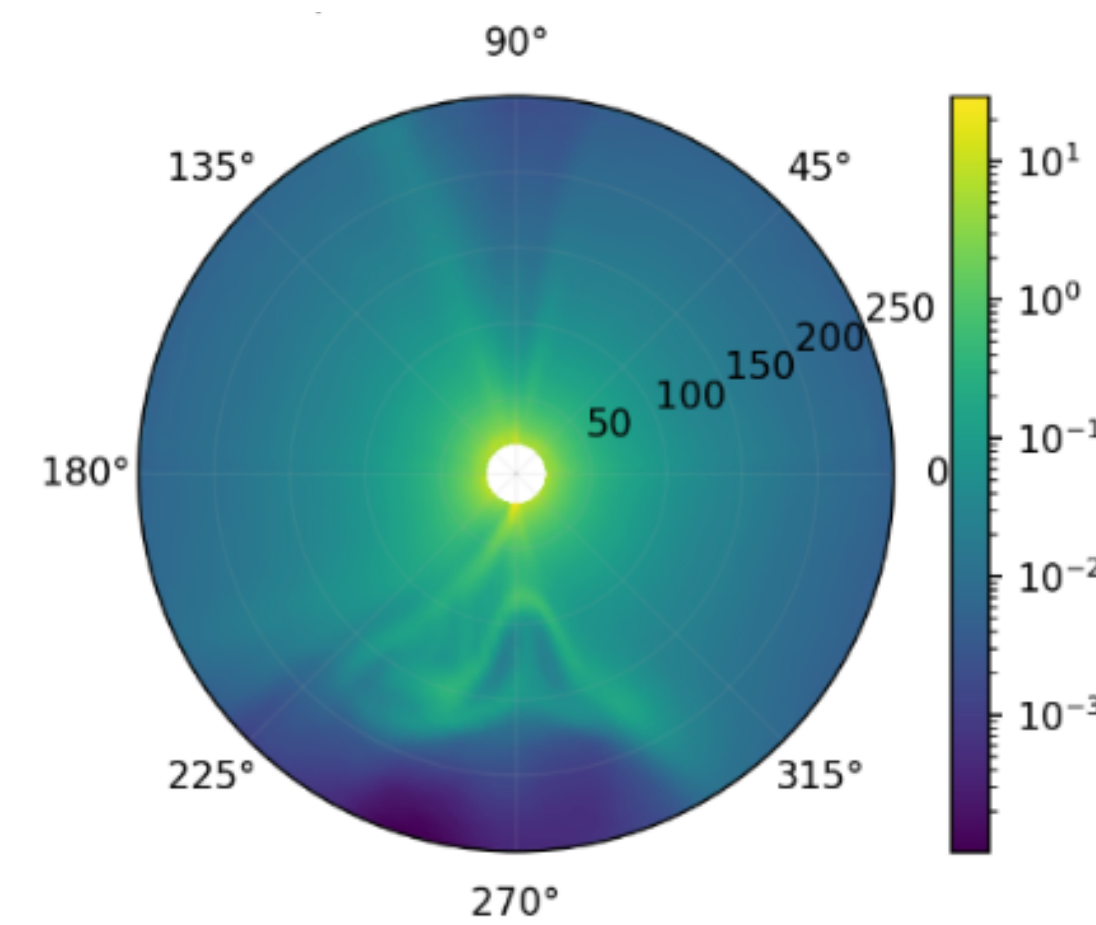
Pressure

MHD

ML

Magnetic Field

MHD



Summary

1. WSA+

- ◉ 40% improvement over traditional WSA model
- ◉ Python package available: *pip install wsaplus*

2. MHD surrogate

- ◉ Full global MHD solution in seconds
- ◉ Can be extended to other PDE systems

3. Sun-Earth Connection

- ◉ Full differentiable pipeline from Sun to (near) Earth
- ◉ Tune with observations to overcome observational constraints